

LBNE Experiment  
 Sanjib R. Mishra  
 University of South Carolina

**Beam**  $\Rightarrow$

🏹 Measurement of the PMNS Matrix Elements

👉  $\Theta_{13}$    👉  $\delta_{CP}$    👉  $\mathbf{V}$ -Mass Hierarchy   👉 Resolving degeneracies

🏹 Beyond PMNS

👉  $\Theta_{23} = 45^\circ$  ?   👉 CPT Violation ?   👉 High  $\Delta m^2$  Oscillation ?

$\Rightarrow$  *Phenomenon that defies the Zeitgeist*

$\Rightarrow$  Syst. Precision

(Nu -vs- NuBar  $\Leftarrow \delta_{CP}$ )

🏹 The familiar, beautiful neighborhood

👉 X-secs,  $\sin^2(\Theta_w)$ : precision comparable to Colliders?

👉 Sum rules, Isospin Physics (Nu -vs- NuBar  $\Leftarrow \delta_{CP}$ )

👉 Heavy neutrinos

👉 .....

👉 Rewriting the  $\mathbf{V}$  text-book

**non-Beam**  $\Rightarrow$

☀ Proton Decay

☀ Supernovae  $\mathbf{V}$  ☀ Solar & Atmospheric  $\mathbf{V}$  ☀ ...

## LBNE Experiment: Salient Parameters

🐜 *Evolution:* MINOS  $\Rightarrow$  NOvA  $\Rightarrow$  LBNE

🐜 *Source:* 700 kW upgraded to 2.3 MW (Project-X)

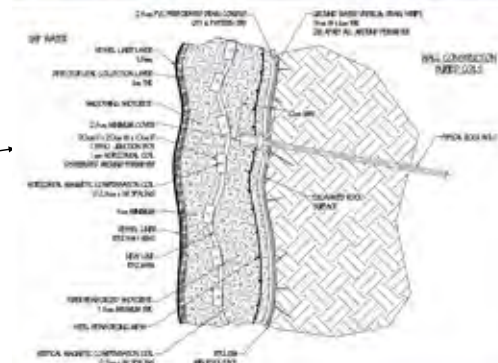
🐜 *Distance:* 1300 km from the  $\nu$  source (Fermilab  $\Rightarrow$  Homestake, SD)  $\Leftarrow$  Mass Hierarchy

🐜 *Beam:* Maximize intensity at the 1st (2.5 GeV) and the 2nd. (0.75 GeV) Maxima  $\Leftarrow \delta_{CP}$

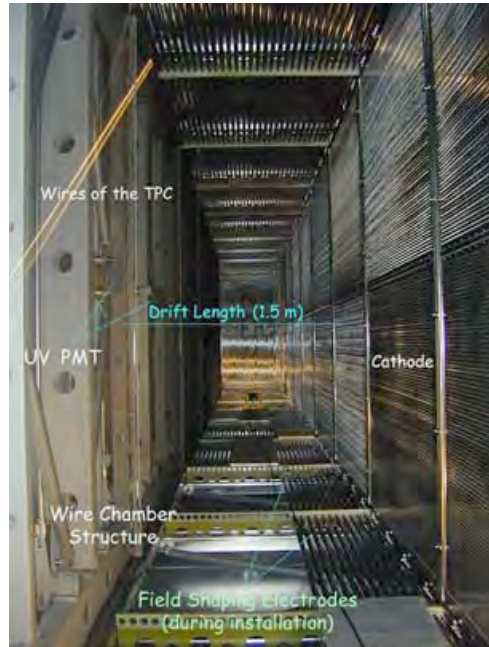
🐜 *Far Detector:* 200 kT Water Cerenkov or 34 kT Liquid Argon

🐜 *Depth:* 4850 ft for WC (possibly LAr)

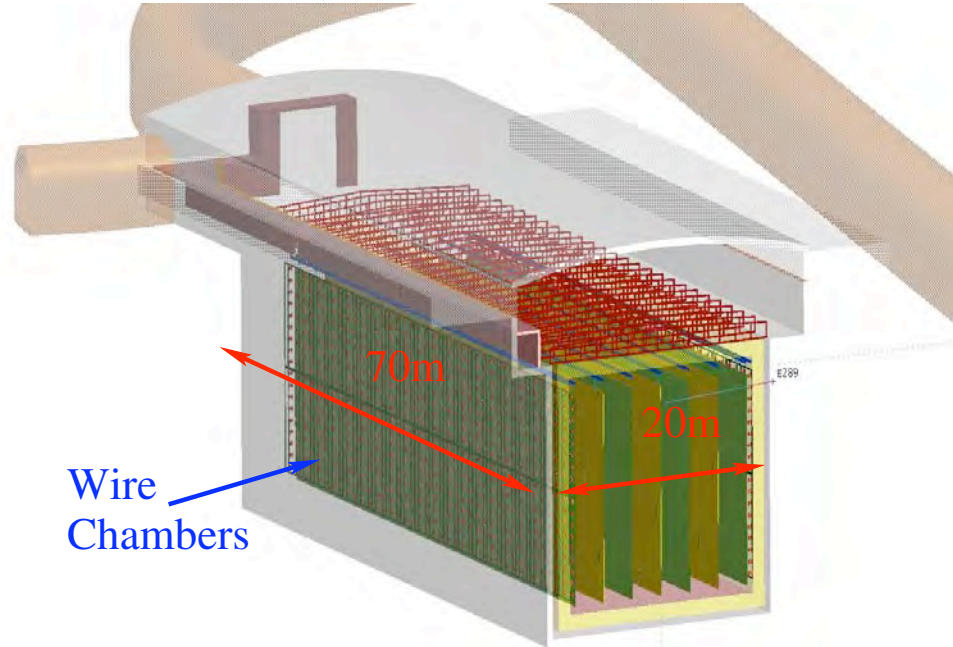
🐜 Comprehensive program to conduct precision measurements at the Near Detector, including  $\nu$ -H<sub>2</sub>O/Ar interactions  $\Leftarrow$  PMNS & non-standard

[illegible]3

**ICARUS : 0.6kt**



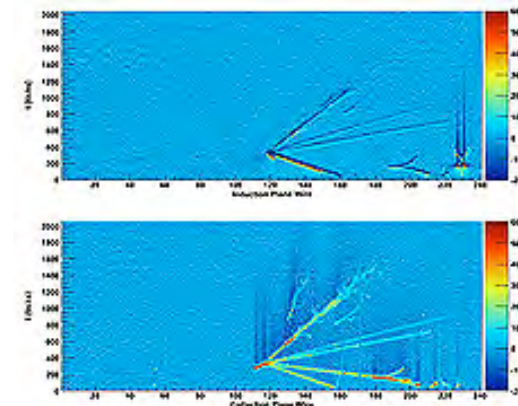
**LBNE LAr : 34 kt module X 2**



**ArgoNeuT (175 litre) prototype in the NuMI beam →**

**High efficiency and purity**

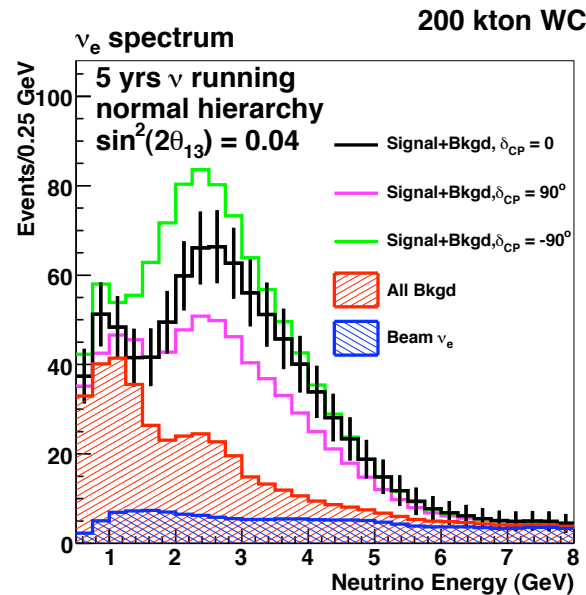
**Requires 60 × scale-up - challenging.**



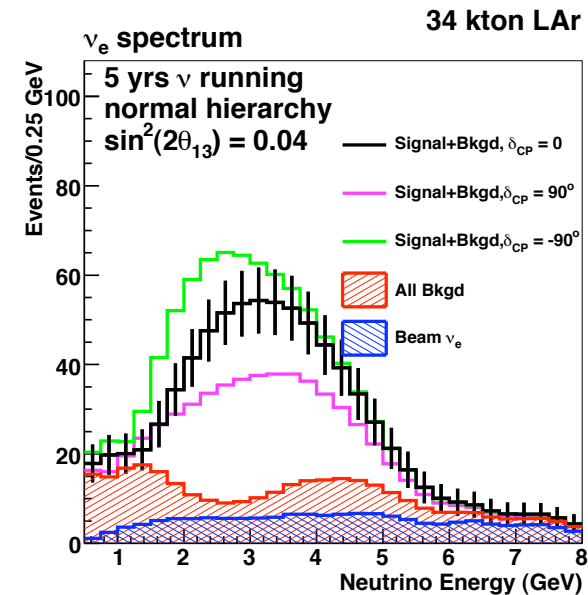
LAR

On-axis wide-band beam (NuMI focusing). Water Cerenkov response is based on the SuperK MC. LAr is modeled as a near-perfect detector. Exposure is 3.5 MW. yr  $\nu$  with  $\sin^2 2\theta_{13} = 0.04$ :

### 200 kt WCD



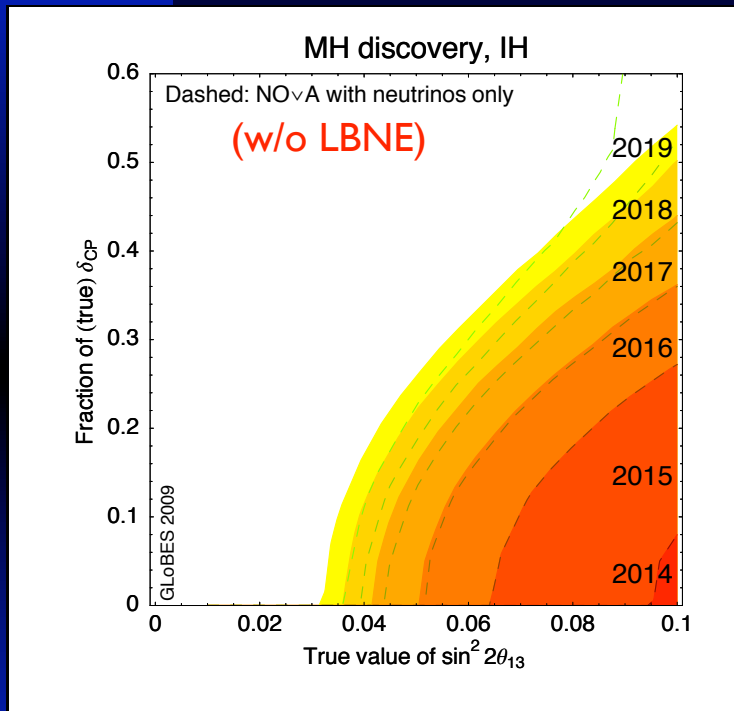
### 34 kt LAr



Interaction rates per 100kT.MW.yr ( $0.5 < E_\nu < 20$  GeV)

$\nu_\mu$ CC	$\nu_\mu$ CC osc	$\nu_e$ CC beam	$\nu_\mu \rightarrow \nu_e$ CC	$\nu_\mu \rightarrow \nu_\tau$ CC
29K	11K	260	560	140

# Mass hierarchy



PH, M. Lindner, T. Schwetz, W. Winter,  
JHEP 11 044 (2009), arXiv:0907.1896.

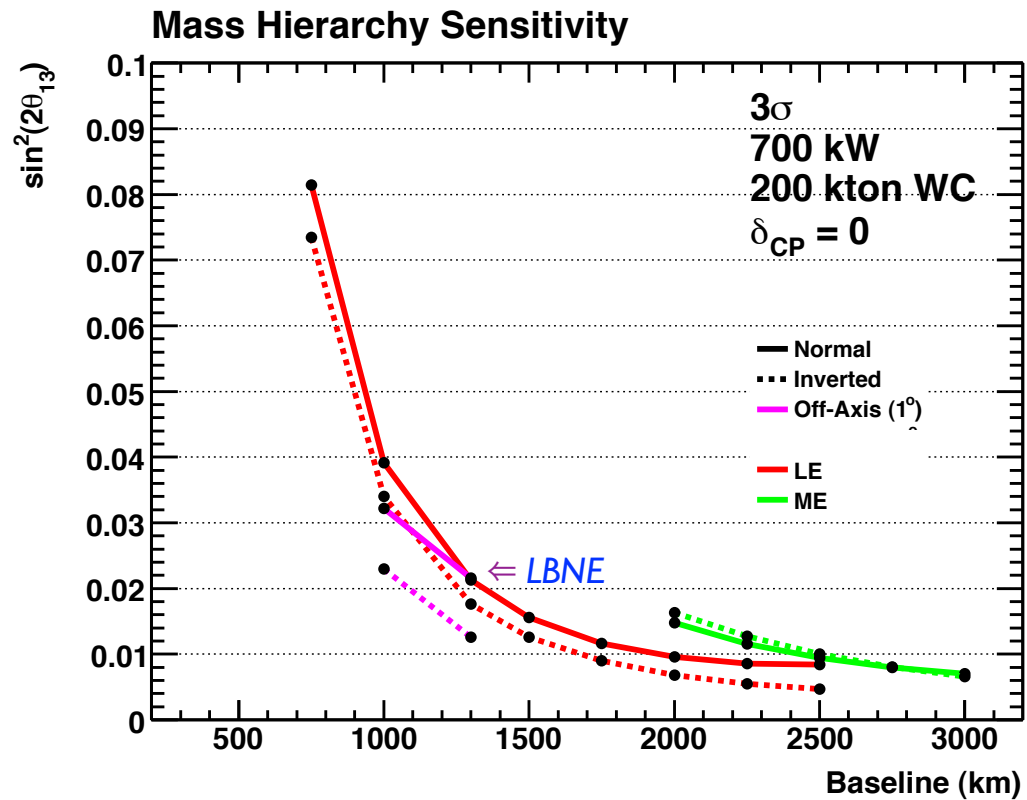
90% CL, combines T2K, NO $\nu$ A, Daya Bay, Double Chooz and RENO At this CL MINOS and T2K have discovered  $\theta_{13} \neq 0$ !

At  $3\sigma$  this plot would be essentially empty!

P. Huber – VT-CNP – p. 3

⇒ Given the  $\Delta m^2_{13}$  and terrestrial medium ...

⇒ Larger Flight Distance implies better MH sensitivity





## Larger $\Theta_{13}$ ? ...& Measurement of $\delta_{CP}$

💡 Larger  $\Theta_{13}$   $\Rightarrow$  Larger number of events:

e.g. for 3.5 MW, 200 kT WVC:

$$\sin^2(2\Theta_{13}) = 0.10 \Rightarrow 1300 \nu_{\mu} \rightarrow \nu_e \text{ events}$$

$$\sin^2(2\Theta_{13}) = 0.01 \Rightarrow 150 \nu_{\mu} \rightarrow \nu_e \text{ events}$$

Measure  $\delta_{CP}$  using:

$$\text{💡 } A(E_\nu) = \frac{P(\nu_{\mu} \rightarrow \nu_e) - P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)}{P(\nu_{\mu} \rightarrow \nu_e) + P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)}$$

💡 Whereas  $P(\nu_{\mu} \rightarrow \nu_e)$  & MH improve with bigger  $\Theta_{13}$ , sensitivity to  $\delta_{CP}$  largely unaffected. As  $\Theta_{13}$  increases the Asymmetry becomes smaller

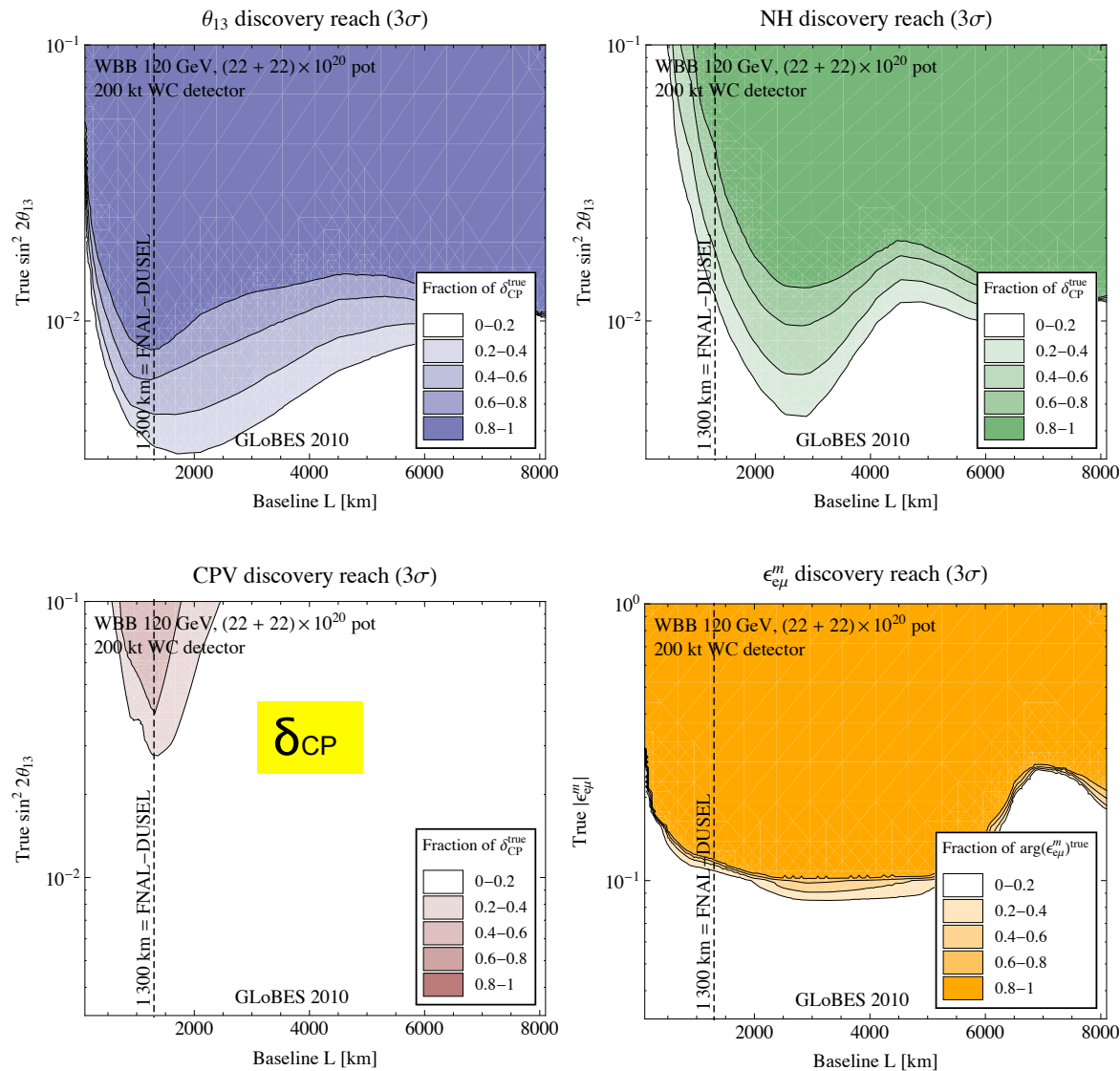
$\Rightarrow$  Longer flight Distance

$\Rightarrow$  Asymmetry at the 2nd. Maximum

$\Rightarrow$  Better control of the Systematics



# Shorter Baseline?



1300 km is optimal for CP at all values of  $\theta_{13}$

The other measurements want a longer baseline

PH, J. Kopp, JHEP 1103:013,2011, ar-Xiv:1010.3706.

## $\delta_{CP}$ vs Asymmetry at the 2nd. Osci. Max ( $E\nu \approx 0.75$ GeV)

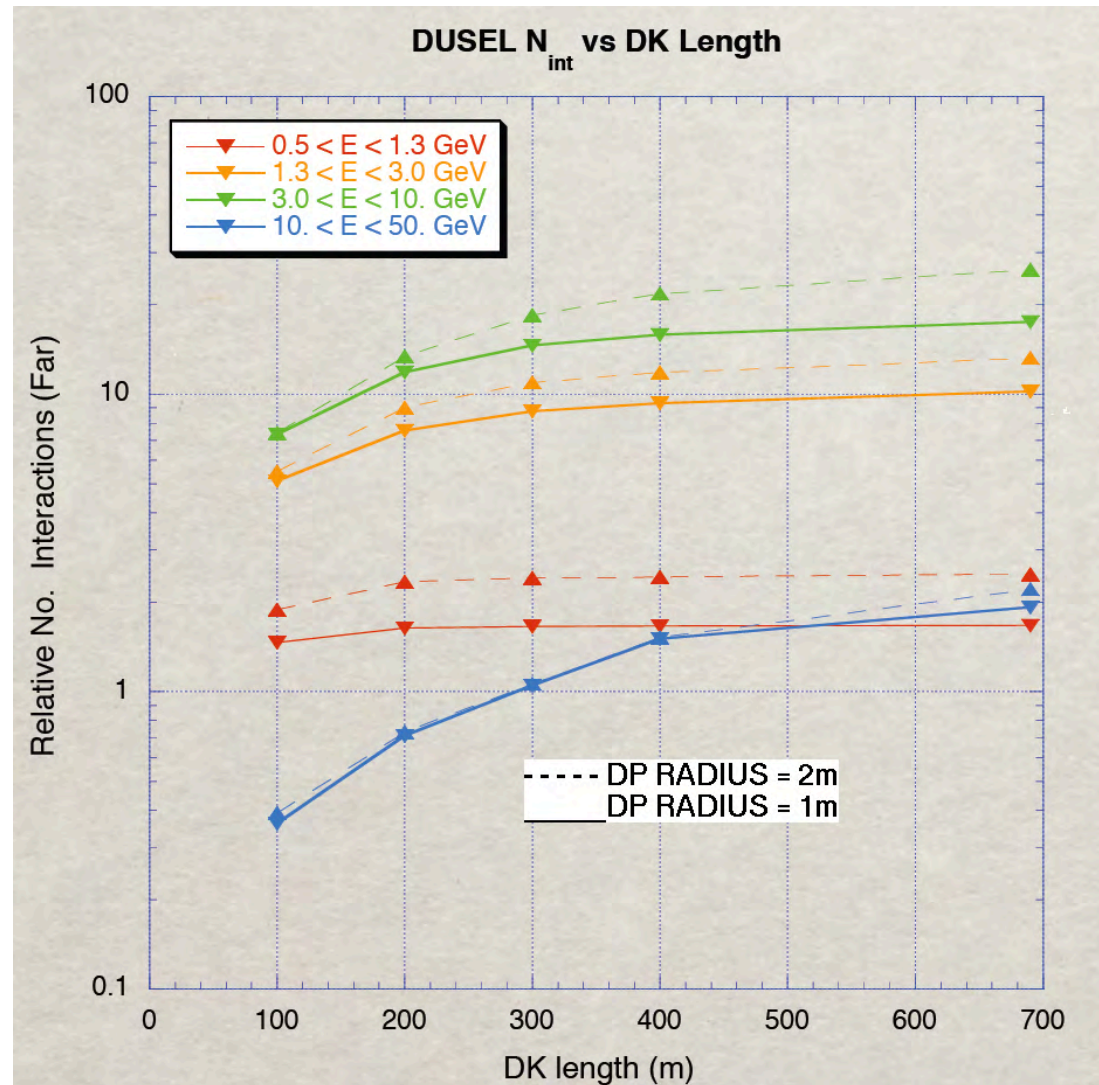
For  $\sin^2(2\theta_{13}) \geq 0.02$ ,  $A(2nd.Max)$  is approx.  $2 \cdot A(1st.Max)$

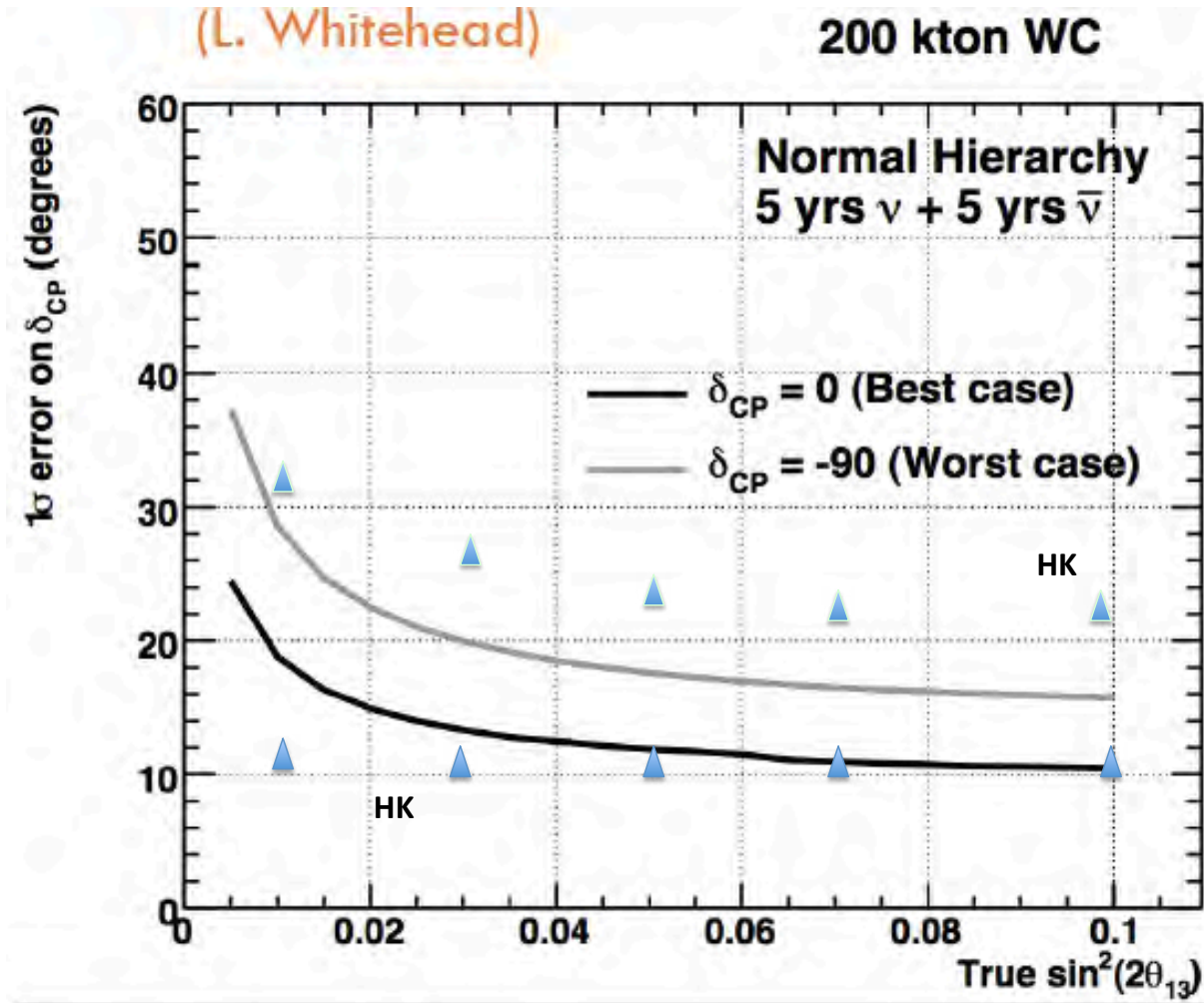
$\Rightarrow$  Maximize the yield at  $E\nu \approx 0.75$  GeV

Note: Unless a new decay pipe is installed for NuMI, the flux will suffer from the small diameter beam pipe of the current setup.

This is true for both first and second oscillation

Need to keep this In mind.

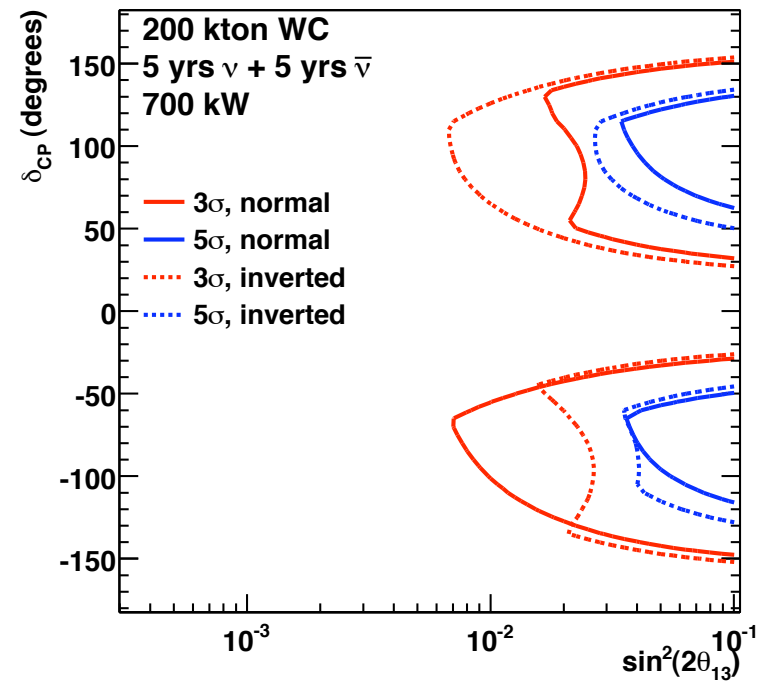
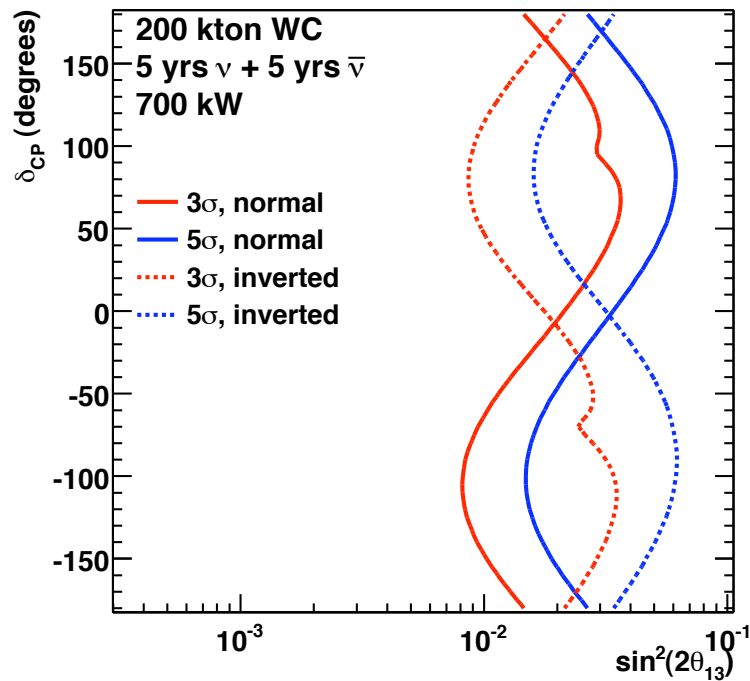




200 kt WCD detector and 5 yrs of  $\nu$  + 5 yrs of  $\bar{\nu}$  running with 700kW:

Mass Hierarchy

CP Violation

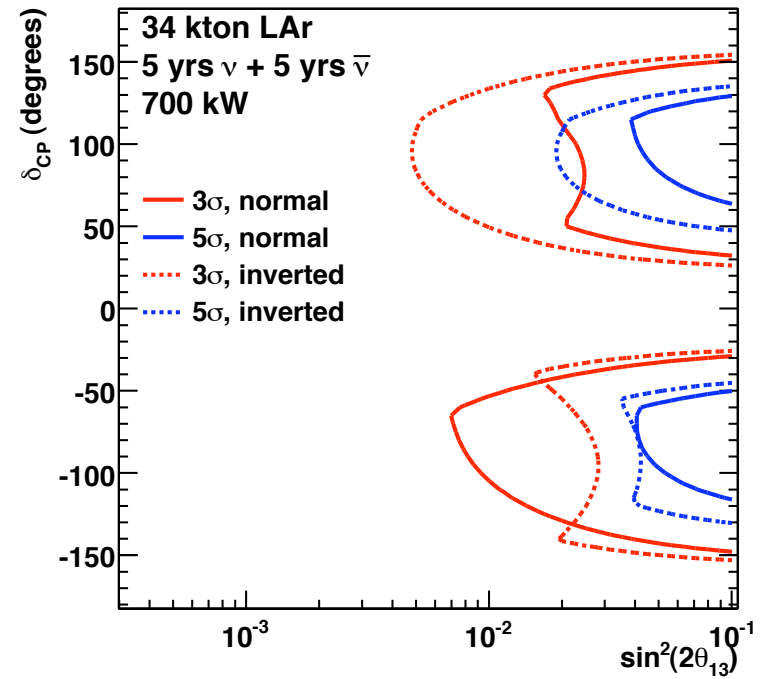
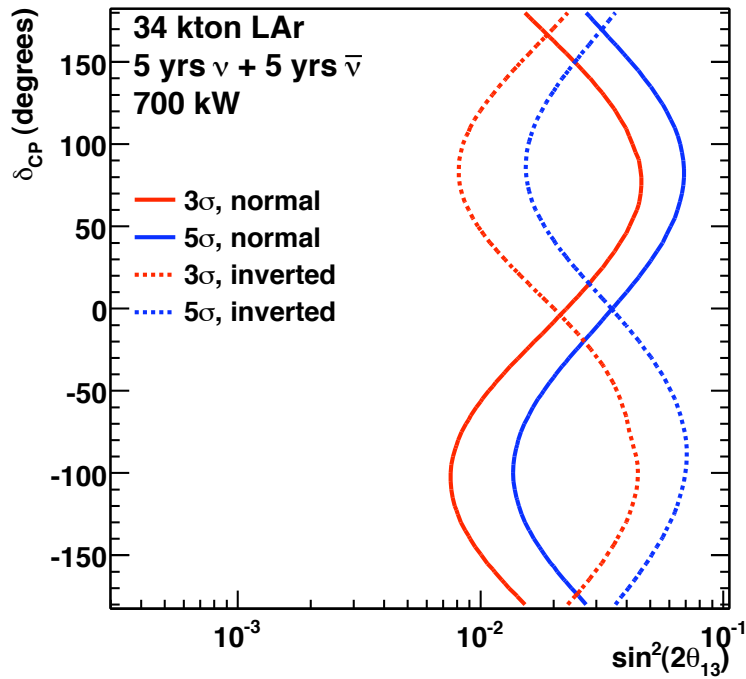


CPV and MH in LBNE 200 kT WC

34 kt LAr detector and 5 yrs of  $\nu$  + 5 yrs of  $\bar{\nu}$  running with 700kW:

Mass Hierarchy

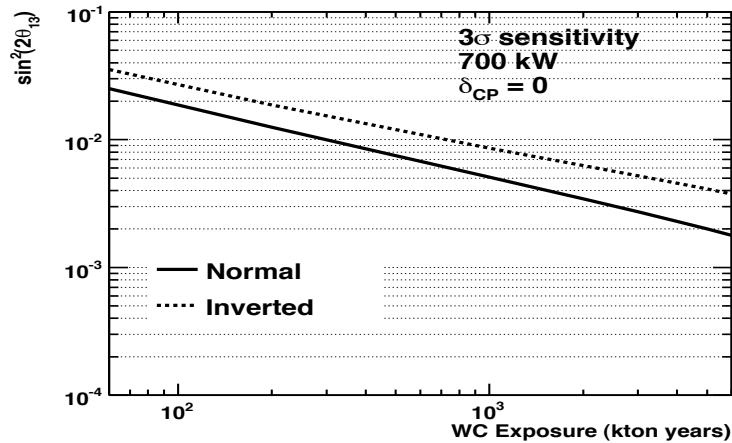
CP Violation



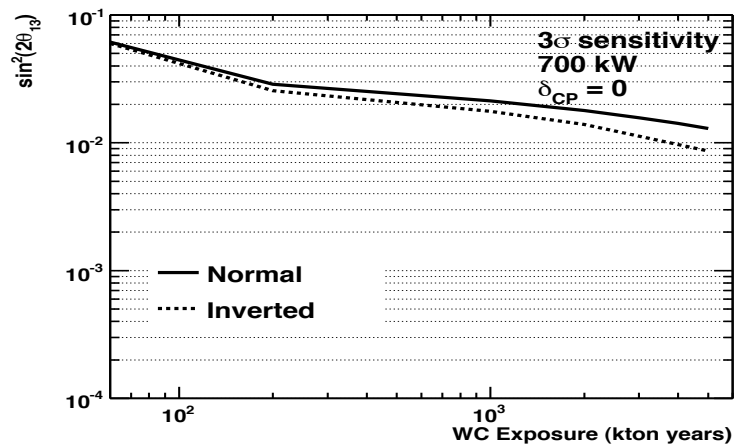
CPV and MH in LBNE 34 kT LAr

## Sensitivity to the Trinity

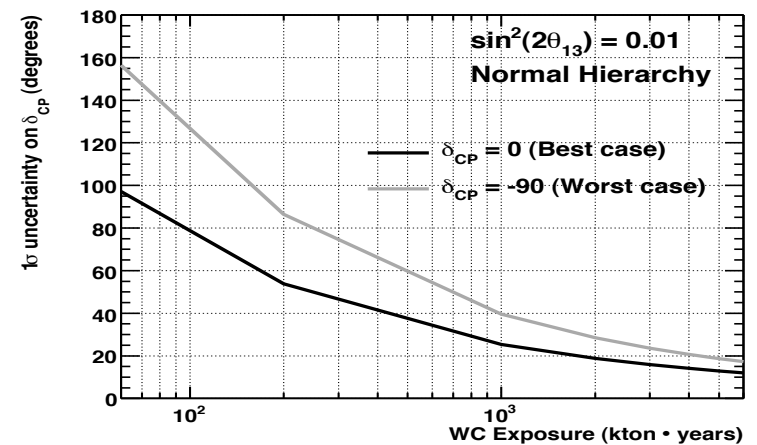
### $\sin^2 2\theta_{13}$ Sensitivity at $3\sigma$



### MH Sensitivity at $3\sigma$



### Resolution of $\delta_{CP}$



### Sensitivity to MH and CPV at $3\sigma$ for $\sin^2 2\theta_{13} \geq 0.01$

## Systematic Errors in $\nu$ Oscillation Experiments

Experiment	$L(\text{km})/E(\text{GeV})$	ND?	Bkg [Expt. Dependent]	Error	$\sin^2(2\theta)$
🐛 BNL-E734	0.1/1.0	No	{418- <b>Ve</b> , 235- <b>NC</b> }	From Data	$3.4 \times 10^{-3}$
🐛 NOMAD {Very fine grain detector}	0.4/25	No	{6000- <b>Ve</b> , 0- <b>NC</b> }	Small	$1.3 \times 10^{-3}$
🐛 MiniBOONE	0.4/0.7	No	{249- <b>Ve</b> , 137- <b>NC</b> }	9%	$\sim 3 \times 10^{-3}$
🐛 MINOS	735/3	Yes	{5- <b>Ve</b> , 44- <b>NC</b> }	5.5%	$\sim 1 \times 10^{-1}$

⇒ Goal for LBNE is <5% systematic error  
Try to achieve it in 'Differential-bins' for NC and CC



## Reinventing the Near Detector

- ◆ Use of “*identical*” *small detector* at the near site is *insufficient* for future  $\mathcal{L}\mathcal{B}\mathcal{L}$  experiments:

- $\Phi^{\nu,\bar{\nu}}(E_\nu, \theta_\nu)$  different at Near & Far sites;
- Impossible to have “*identical*” detectors, for  $\mathcal{O}(100\text{kt})$ , at the projected luminosities;
- Different compositions of event samples ( $\nu_\mu, \bar{\nu}_\mu, \nu_e$ , NC, CC)

⇒ Coarse resolution dictated by  $\mathcal{O}(100\text{kt})$  and different flux at Near-vs-Far tell us that the *Identical Near Detector* concept is insufficient

- ◆ Need a *high resolution detector* at the Near-Site to measure systematics affecting the Far-detector:

- $\nu_\mu, \bar{\nu}_\mu, \boxed{\nu_e}, \boxed{\bar{\nu}_e}$  content vs.  $E_\nu$  and  $\theta_\nu$ ;
- $\nu$ -induced  $\pi^\pm/K^\pm/p/\pi^0$  in CC and NC interactions;
- Quantitative determination of  $E_\nu$  absolute energy scale;
- Measurement of detailed event topologies in CC & NC.

⇒ Provide an ‘Event-Generator’ measurement for  $\mathcal{L}\mathcal{B}\mathcal{L}\nu$

☞ Measure over the full range of FD  
☞ Background to the  $\mathbf{V}(\text{Bar})e/\mu$ -Appearance  
☞  $\mathbf{V}$  -vs0  $\mathbf{V}(\text{Bar})$  Interactions

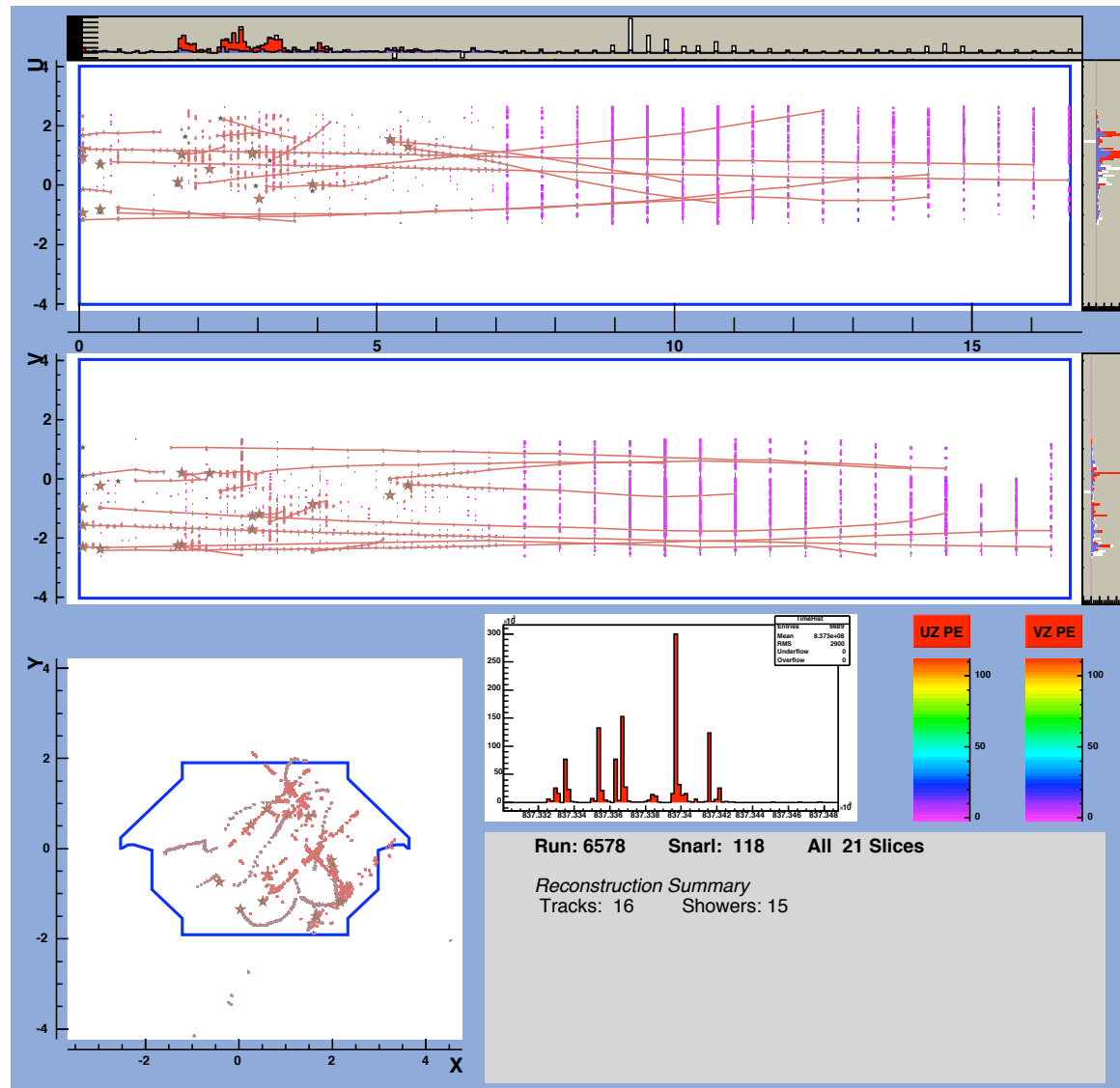
- ◆ High Resolution near detectors at future  $\mathcal{L}\mathcal{B}\mathcal{L}$  facilities are natural heirs to the *precision neutrino scattering programme*

Can they achieve sufficient precision to complement the Colliders?

## Events/Spill in MINOS-ND

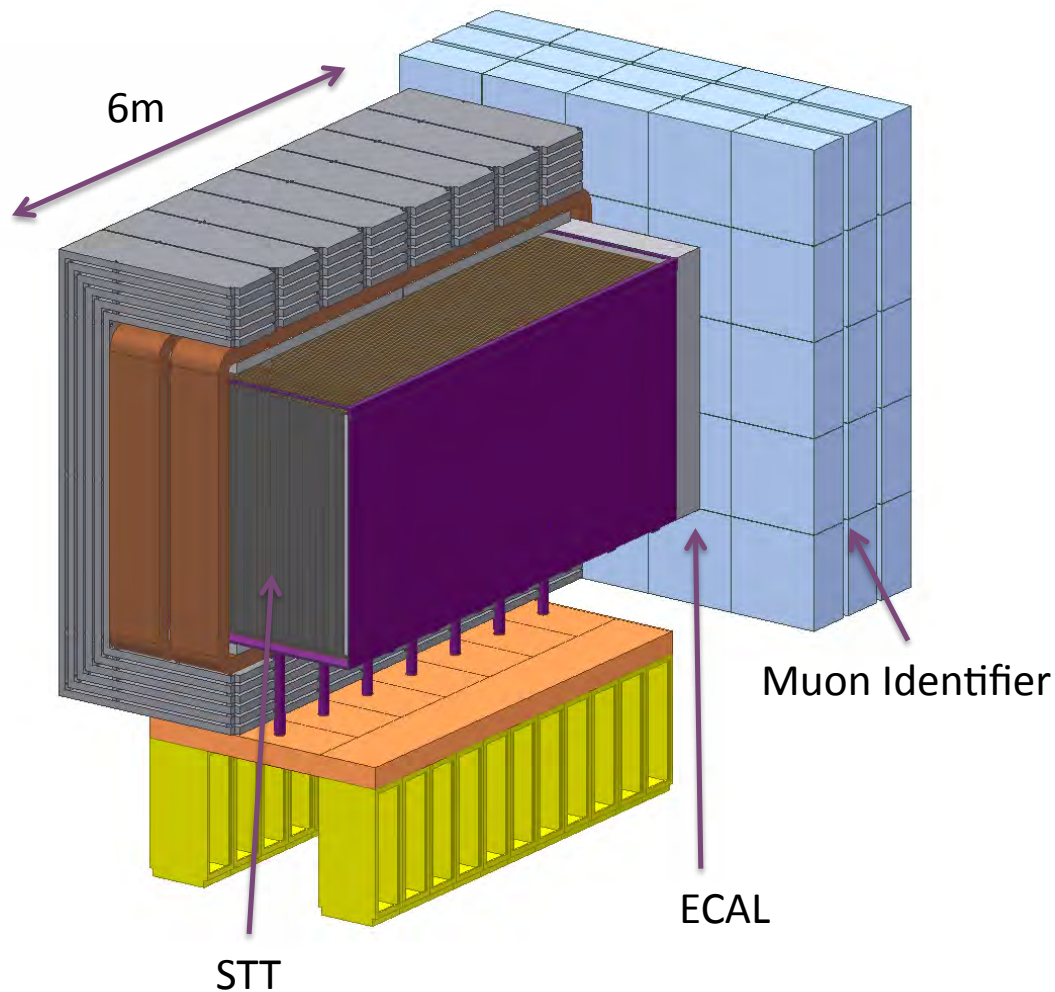
$\approx 2 \times 10^{13}$  PoT/Spill in MINOS(NuMI)

(Juxtapose against that expected from  $3 \times 10^{14}$  PoT/Spill in ProjectX)



10 / 24

# Straw Tube Tracker (STT)



📌 Best performance of the 4-options

📌 3.5m x 3.5m x 7m STT (7 tons;  $\rho \approx 0.1 \text{ gm/cm}^3$ )

4 $\pi$ -ECAL

Dipole-Field (0.4T)

$\mu$ -Detector (RPC) in Dipole and Downstream

Transition Radiation  $\Rightarrow$  e-/e+ ID  $\Rightarrow$   $\gamma$

dE/dx  $\Rightarrow$  Proton,  $\pi^+/-$ ,  $K^+/-$

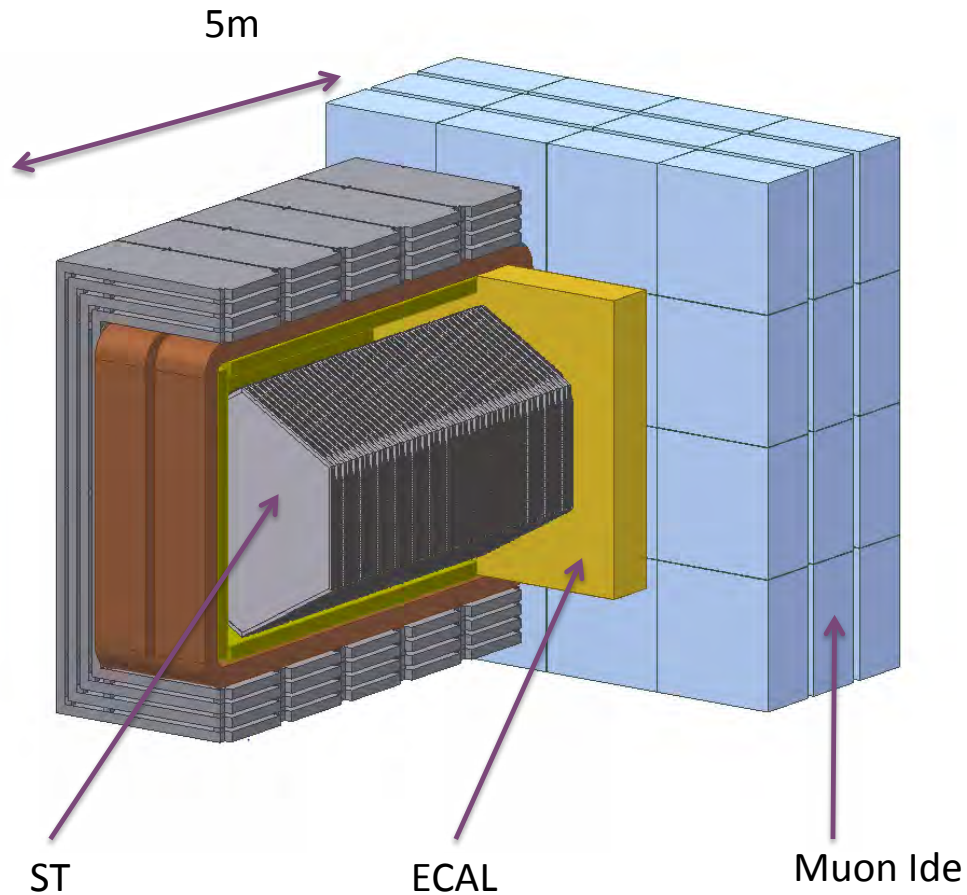
Magnet/Muon Detector  $\Rightarrow$   $\mu^+/-$

📌 H<sub>2</sub>O & D<sub>2</sub>O Targets ( $\approx$  x5 FD-Stat)  $\Rightarrow$  WC-FD

{QE-Proton ID  $\Rightarrow$  Absolute Flux measurement}

📌 Pressurized Ar-target ( $\approx$  x5 FD-Stat)  $\Rightarrow$  LAr-FD

# Scintillator Tracker (ST)



3m x 3m x 5m Sci-Tracker (7 tons;  $\rho \approx 1 \text{ gm/cm}^3$ )

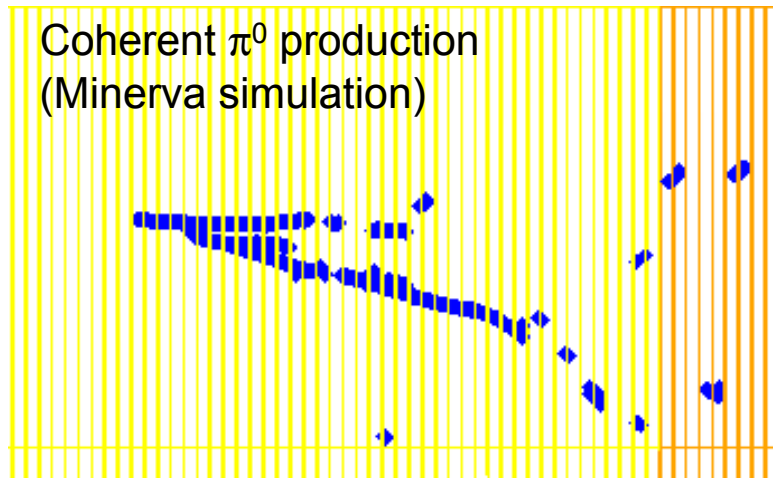
4 $\pi$ -ECAL

Dipole-Field (0.4T)

$\mu$ -Detector (RPC) in Dipole and Downstream

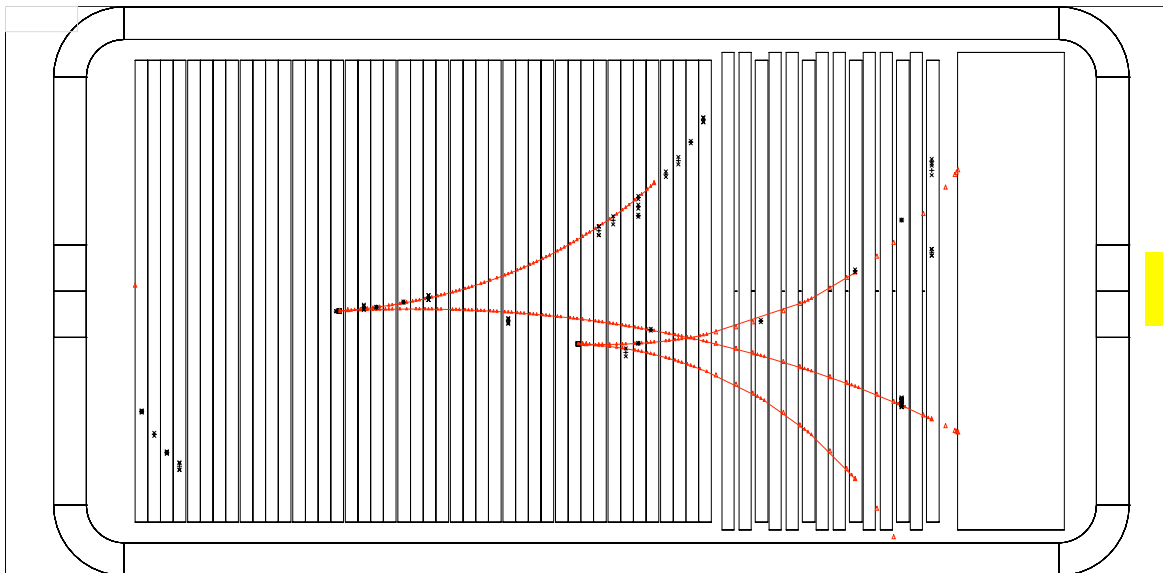
H<sub>2</sub>O Target ( $\approx \times 5$  FD-Stat)  $\Rightarrow$  WC-FD

## Coh- $\pi^0$



A Question of Resolution...

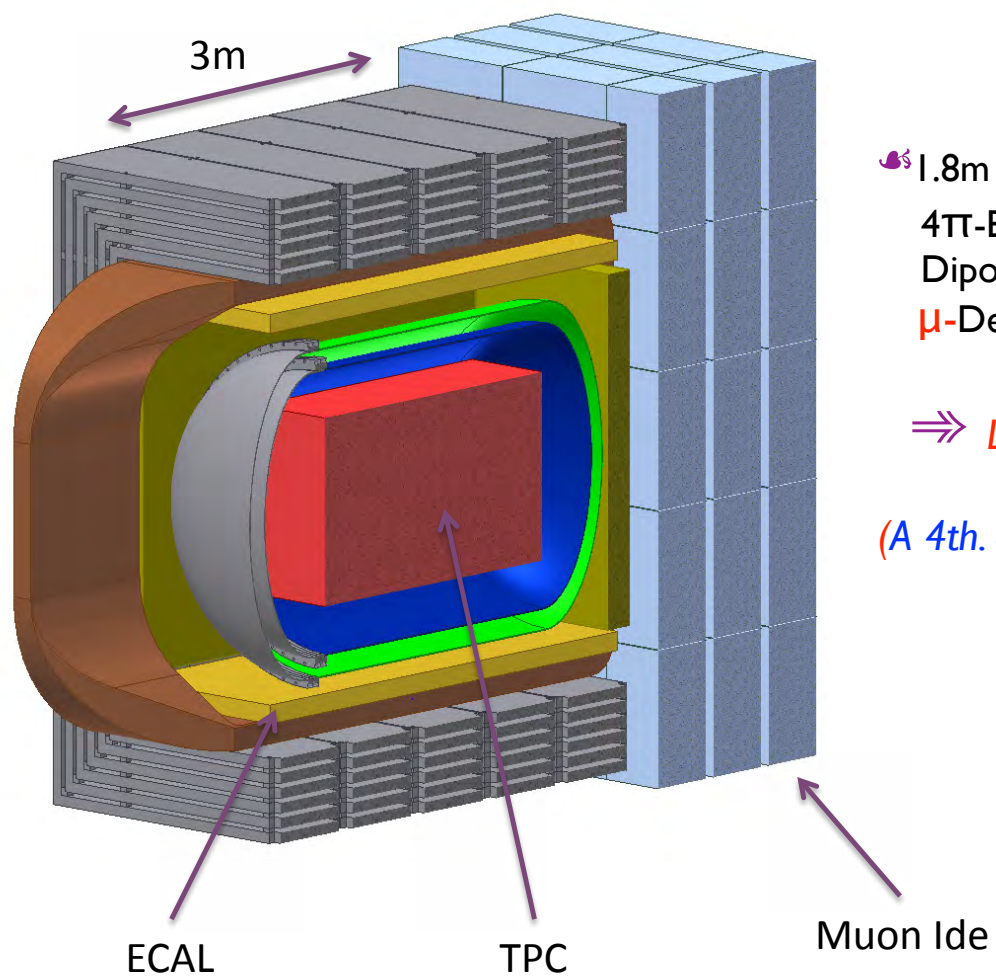
## NOMAD DATA



⇐ x12 more hits in STT

(Hits shown by 'x' are not used in the track-fit)

# LAr TPC Tracker (TPCT)



1.8m x 1.8m x 3m LAr (13 tons)

4T-ECAL

Dipole-Field (0.4T)

$\mu$ -Detector (RPC) in Dipole and Downstream

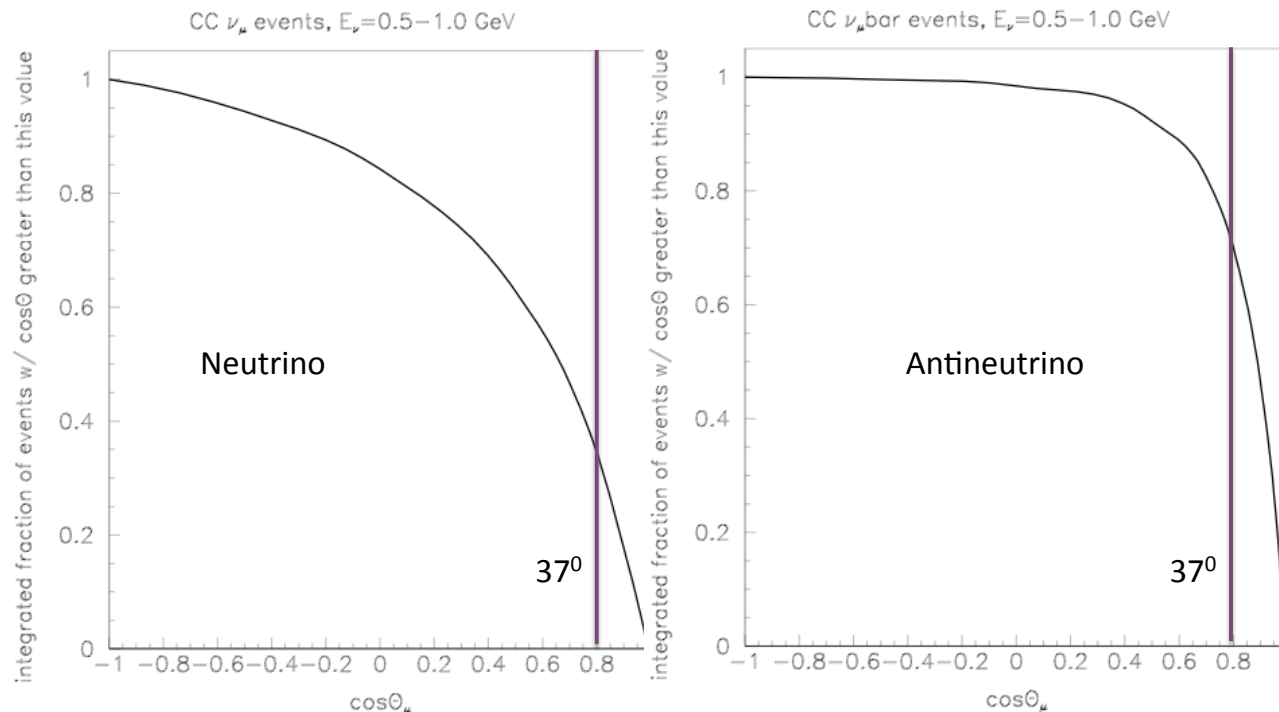
$\Rightarrow$  LAr-FD

*(A 4th. option with larger LAr with membrane cryostat B-Field)*

## Why Tracker (ECAL/ $\mu$ -Detector) within a B-Field?

- ☞ Constrain  $E_\nu$ -scale
  - ☞ ND must measure the full range of  $E_\nu$  &  $\theta_\nu$  else the sensitivity of FD will be compromised
  - ☞ In  $0.5 \leq E_\nu \leq 1$  GeV, the Acceptance  $\approx 35\%$  for  $\theta_\mu \leq 37^\circ$   
In  $2.0 \leq E_\nu \leq 3$  GeV, the Acceptance  $\approx 75\%$  for  $\theta_\mu \leq 37^\circ$
  - ☞ For LBNE, the Maximal sensitivity for  $\delta_{CP}$  is  $E_\nu \approx 1.5$  GeV  
*Measure differences in  $\nu$  & Anti- $\nu$  interactions which might fake a “ $\delta_{CP}$ ”*
  - ☞ STT will be able to distinguish  $\mu^-/\mu^+$  down 0 ~0.3 GeV
- ⇒ ND must measure and ID leptons (at least  $\mu$ ) emerging at large angles;

0.5-1 GeV



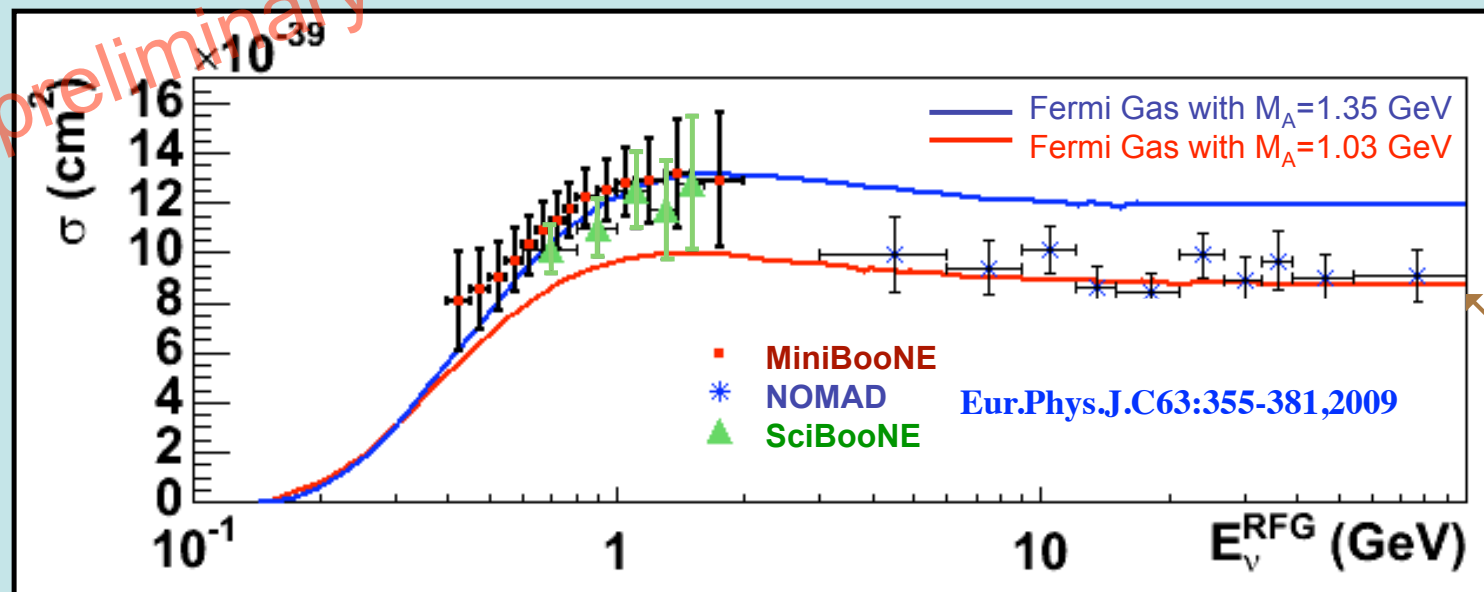


## Why track protons?

- 💡 Precision determination of  $\nu_\mu$ -QE requires proton-tracking.
  - ⇒ QE in H<sub>2</sub>O & D<sub>2</sub>O will provide an Absolute-Flux measurement:
    - Need proton-tracking & resolution to point to the H<sub>2</sub>O & D<sub>2</sub>O vertex
  - ⇒ ( $\mu^-$ , p) provide an *in situ* constraint on the Fermi-motion and hence on the  $E_\nu$ -scale
  - ⇒ QE interactions dominant in Low- $E_\nu$ : Need accurate parametrization of QE [see  $\sigma(QE)$  Fig.]
- 💡 STT option will have a large proton sample from  $\Lambda \rightarrow p \pi^-$
- 💡 If an ND is able to accurately measure proton, it will be able to measure the  $\pi^-$  &  $\pi^+$  in NC and CC:
  - the largest source of background to the  $\nu_\mu$  & Anti- $\nu_\mu$  disappearance*
- ⇒ *ND must track & ID QE-protons*

# Quasi-Elastic Scattering

- new, modern measurements of QE  $\sigma$  at these energies (on  $^{12}\text{C}$ )



~ 30% difference between QE  $\sigma$   
measured at low & high E on  $^{12}\text{C}$  ?!

?

## $\nu_\mu$ -QE Sensitivity Calculation

Example of a V-interaction in a high-resolution ND as a calibration of FD

Key is 2-Track ( $\mu$ , p) signature *Proton reconstruction: the critical issue*  
(*dE/dx in but not used in the analysis*)

Use Nomad data/MC as calibration

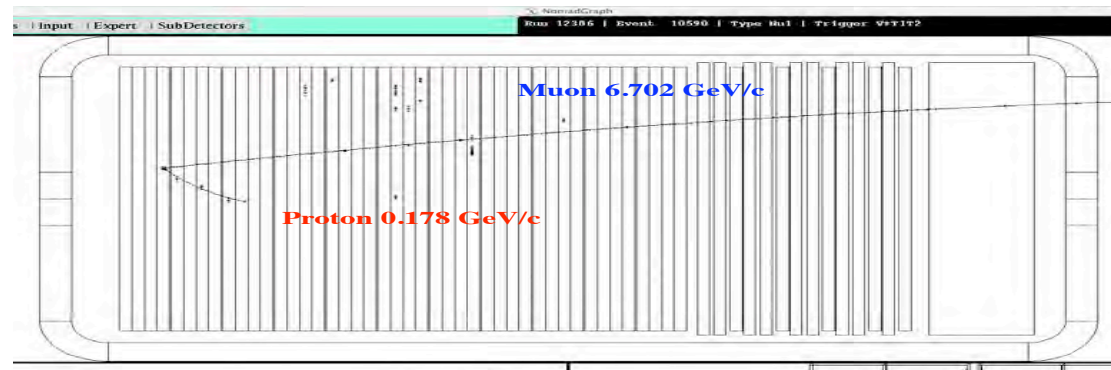


Figure 14: A  $\nu_\mu$ -QE candidate in NOMAD

QE Candidates in NOMAD: STT will have **x6** more points for protons

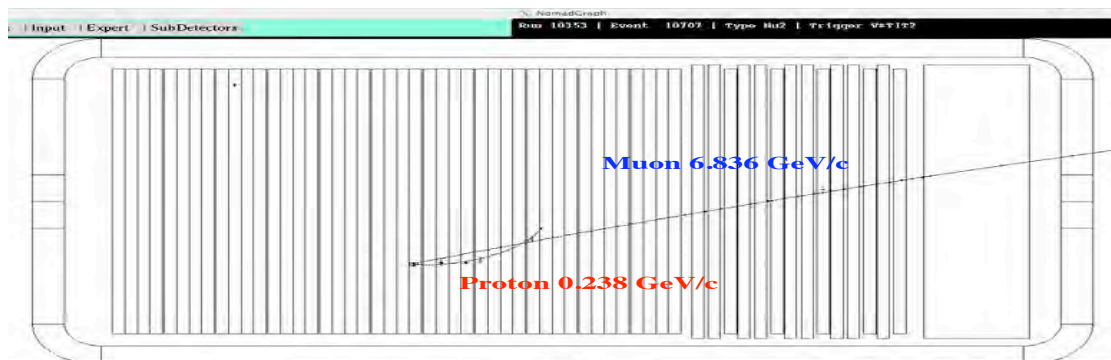
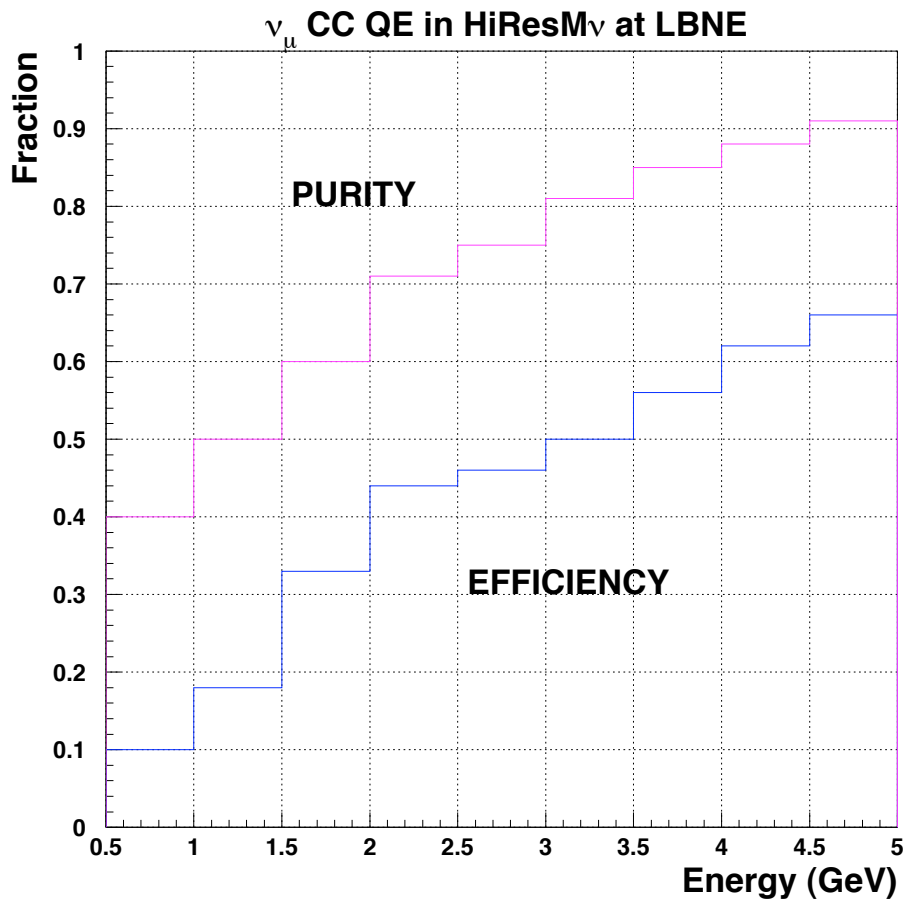


Figure 15: A  $\nu_\mu$ -QE candidate in NOMAD

# RECONSTRUCTION OF CC QUASI-ELASTIC INTERACTIONS



- ♦ Protons easily identified by the large  $dE/dx$  in STT & range  
 $\Rightarrow$  Minimal range to reconstruct  $p$  track parameters 12cm  $\Rightarrow$  250 MeV
- ♦ Analyze BOTH 2-track and 1-track events to constrain FSI, Fermi motion and nuclear effects
- ♦ Use multi-dimensional likelihood functions incorporating the full event kinematics to reject DIS & Res backgrounds  
 $\Rightarrow$  On average  $\varepsilon = 52\%$  and  $\eta = 82\%$  for CC QE at LBNE

## Why measure and ID $e^-$ & $e^+$ ?

💡 Measurement of  $\pi^0$  in NC and CC via  $\gamma \rightarrow e^-e^+$  measured in the tracker

{  $\pi^0$  is the largest background to (anti) $\nu_e$ -appearance }

💡 Measure beam  $\nu_e$  and Anti- $\nu_e$

⇒ Difference between ( $\nu_e$  from  $\mu$ ) & (anti- $\nu_e$  from  $K^0_L$ ) extrapolations to FD from ND

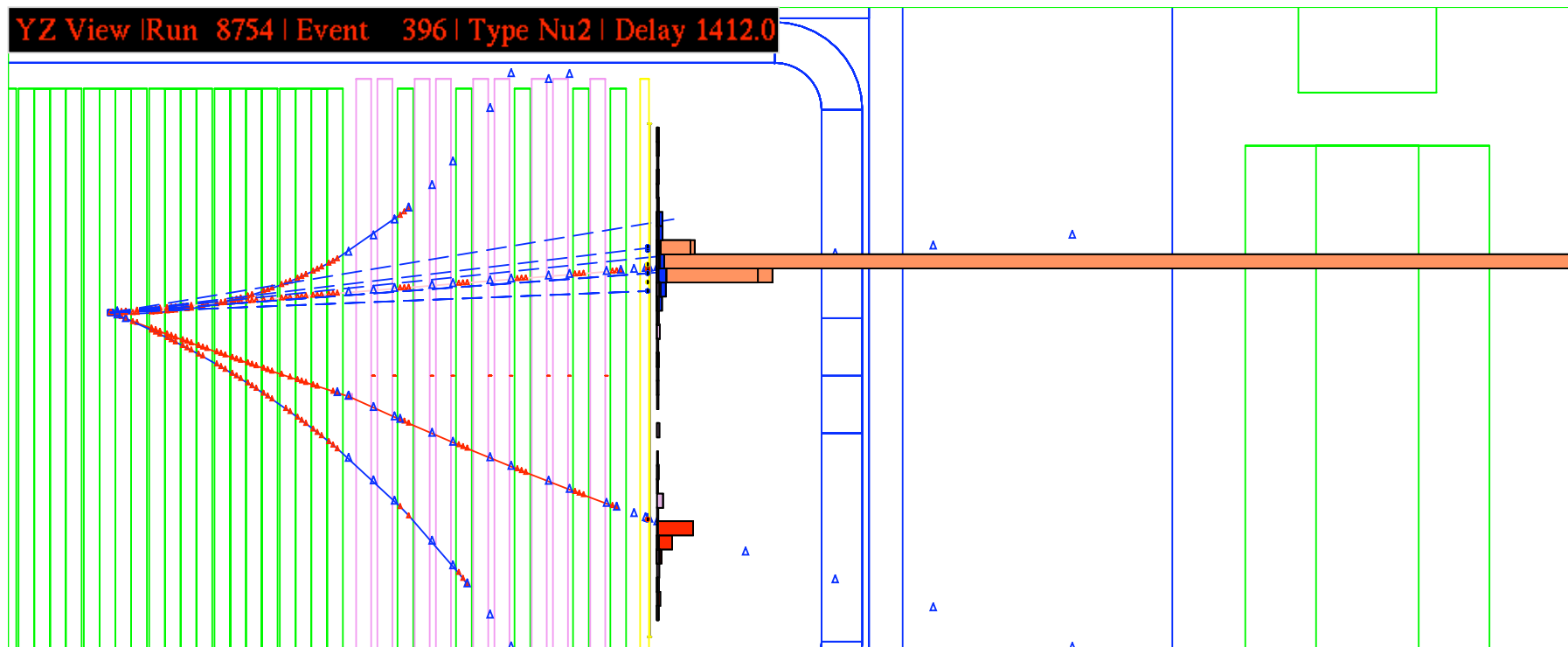
⇒ A must if there are large- $\Delta m^2$  oscillations

💡 Measurement of absolute flux

💡 To discover  $\delta_{CP}$  we ought to ensure that  $\nu_e$  & anti- $\nu_e$  events are as expected

⇒ *ND must measure  $\pi^0$  and  $\nu_e$  & anti- $\nu_e$   $\rightarrow e^-$ -vs-  $e^+$*

## A $\bar{\nu}_e$ CC candidate in NOMAD

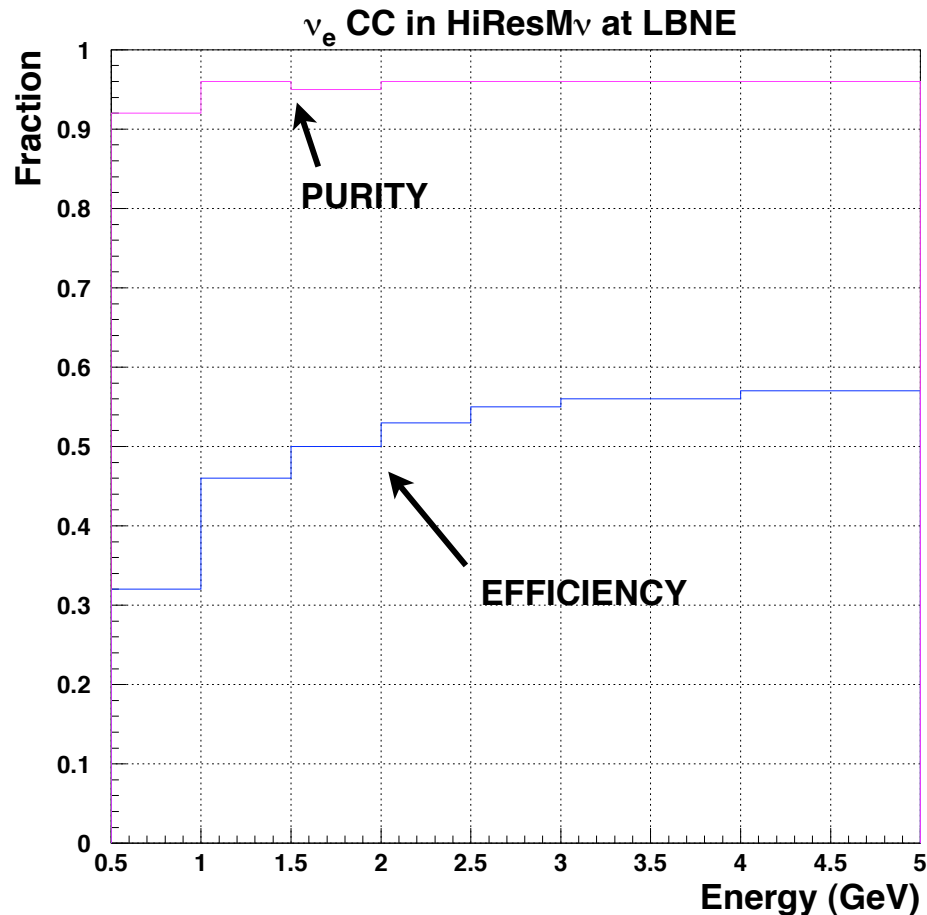


- ☞ x12 higher sampling in STT (HiResMnu)
- ☞ x4 $\pi$  calorimetric and  $\mu$  coverage





## IDENTIFICATION OF $\nu_e$ CC INTERACTIONS



- ♦ The HiResM $\nu$  detector can *distinguish electrons from positrons in STT*  
 $\Rightarrow$  *Reconstruction of the e's as bending tracks NOT showers*
- ♦ Electron identification against charged hadrons from both TR and dE/dx  
 $\Rightarrow$  *TR  $\pi$  rejection of  $10^{-3}$  for  $\varepsilon \sim 90\%$*
- ♦ Use *multi-dimensional likelihood functions* incorporating the full event kinematics to reject non-prompt backgrounds ( $\pi^0$  in  $\nu_\mu$  CC and NC)  
 $\Rightarrow$  *On average  $\varepsilon = 55\%$  and  $\eta = 99\%$  for  $\nu_e$  CC at LBNE*

👉 **VeBar-CC Sensitivity:**

If we keep the **signal efficiency** at **~55%**, then **purity** is about **95%**

# Absolute Flux using $\nu$ -e Elastic NC Scattering

Using the Weak Mixing Angle (0.238) at  $Q \sim 0.1$  GeV (known to  $\leq 1\%$  precision)

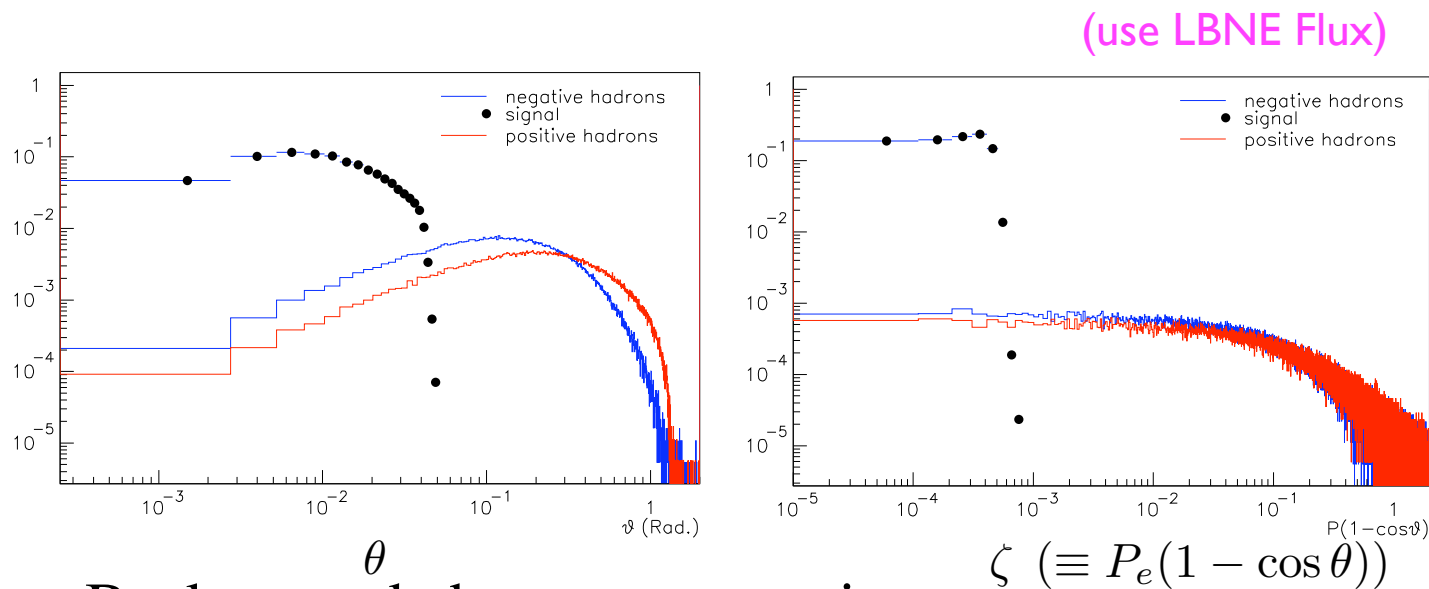
$\Rightarrow \sigma(\nu_e e \text{-NC})$  known  $\Rightarrow$  Absolute- $\phi(\nu_e)$

$\nu_e e \rightarrow$  Signal: Single, forward  $e^-$

Background: NC induced  $\pi^0 \rightarrow \gamma \rightarrow e^-$  ( $e^+$  invisible): charge-symmetric

Two-step Analysis:  $\star$  Electron-ID: TR  $\star$  Kinematic cut:  $\zeta = P_e(1 - \cos\theta_e)$

Simulation of charged hadron background.



Background charge symmetric & benign

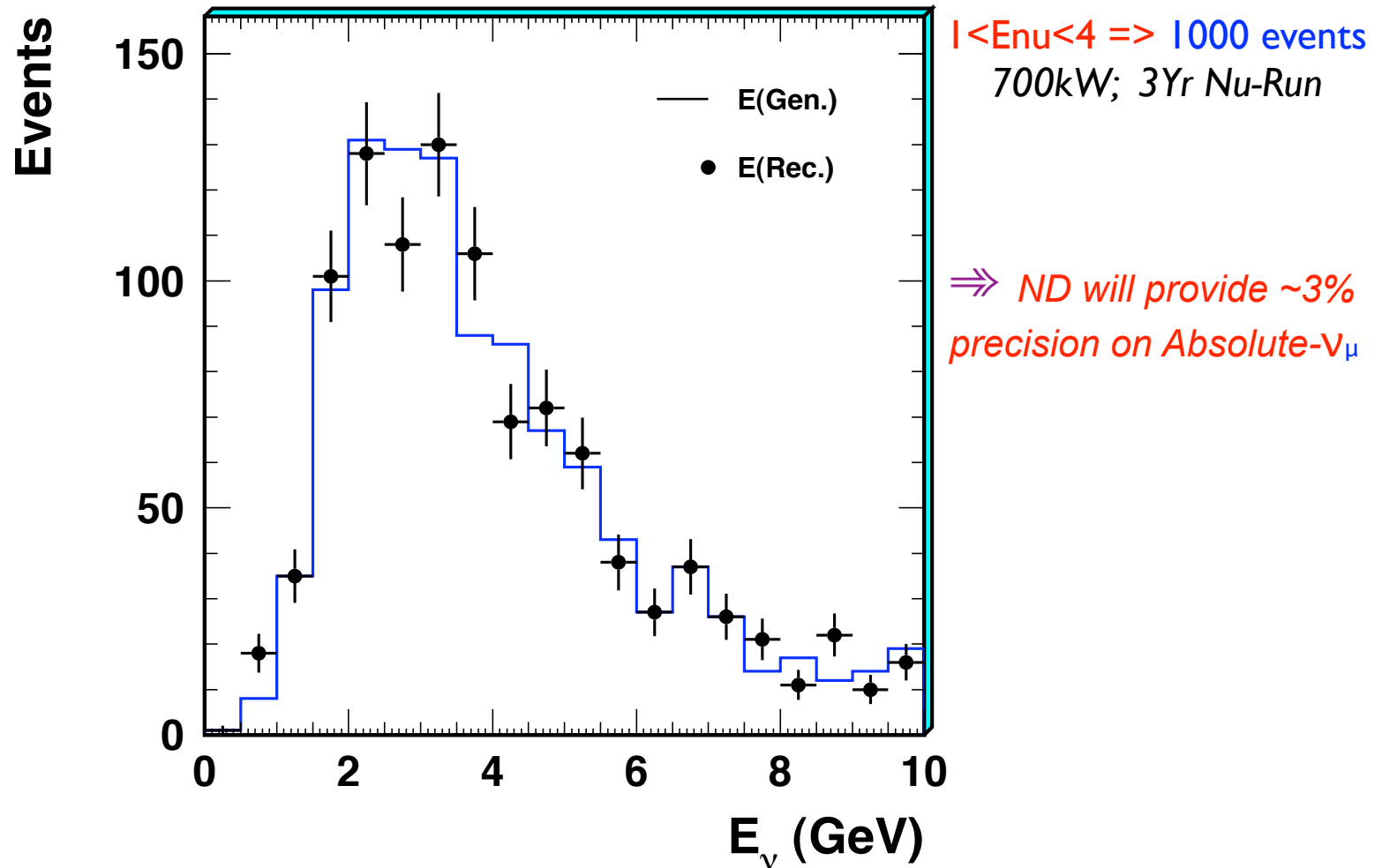
Eff > 64%  
Bkg >  $\leq 10^{*-6}$   $\Leftarrow$  Measured

$\Leftarrow$  Conclusion

## Absolute Flux using $\nu$ -e Elastic Scattering

🔗 Shape of  $E_{\nu}$  using ( $E_e$ ,  $\theta_e$ ):

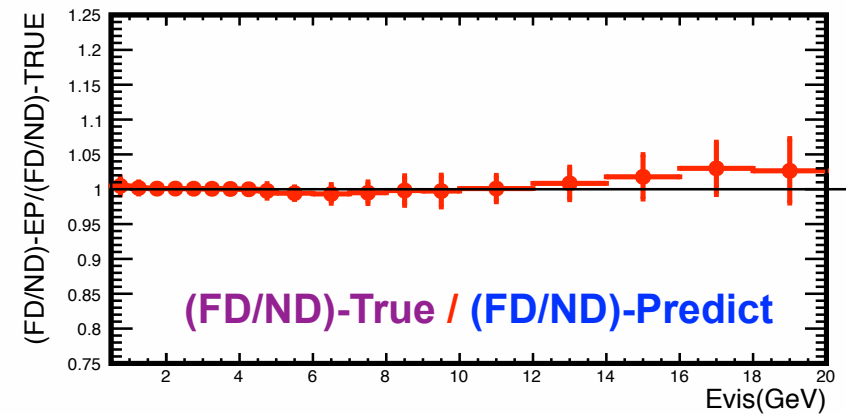
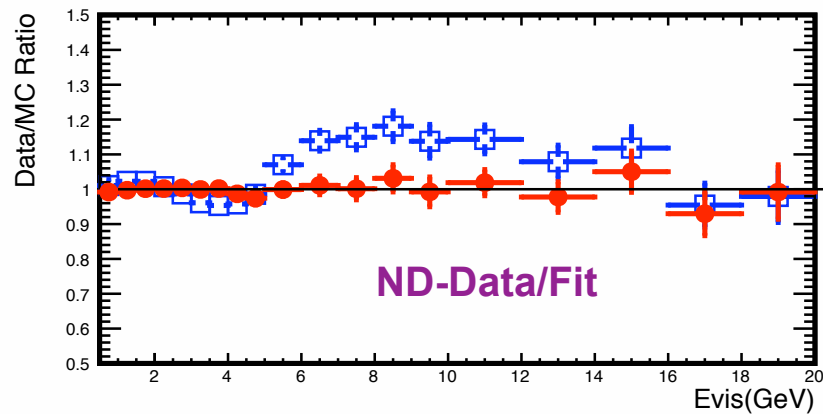
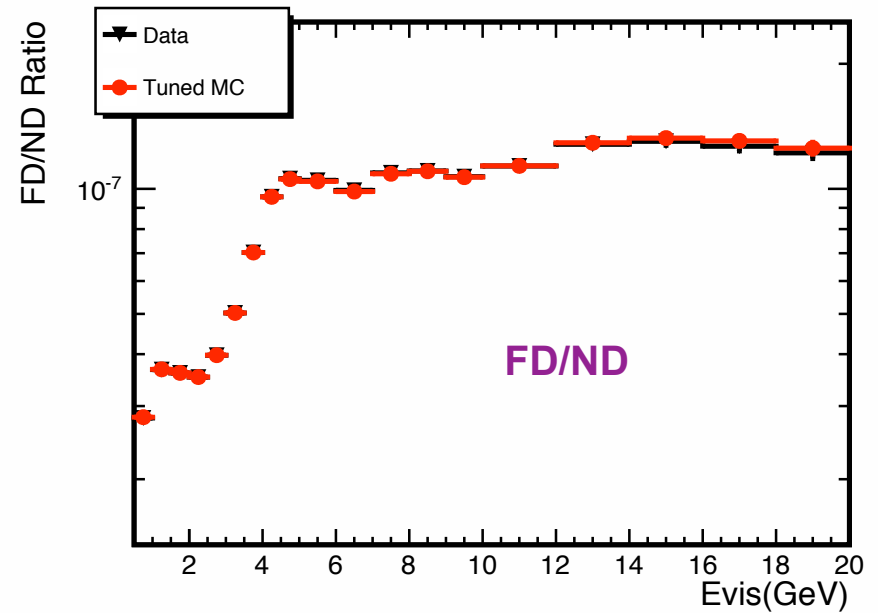
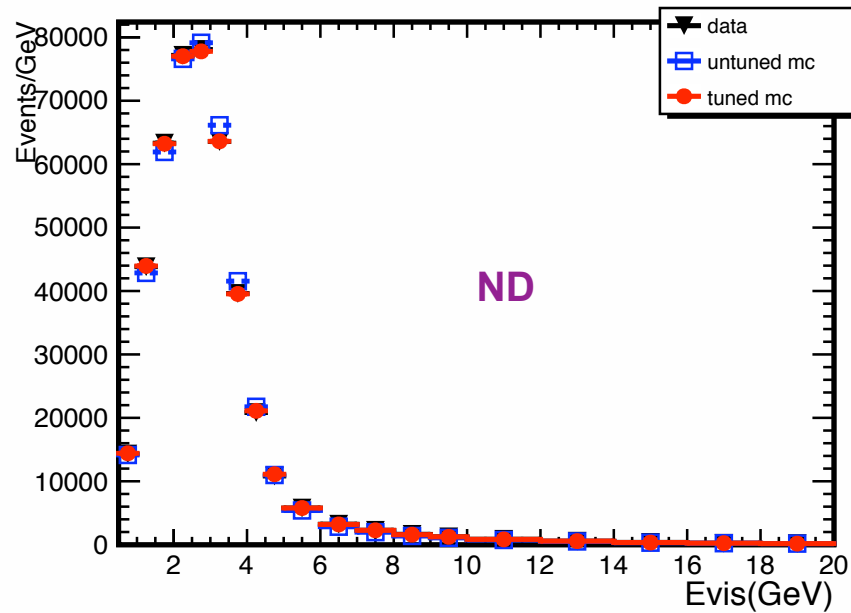
🔗 The precision on relative  $\nu$ -flux (shape) is worse than in that determined using Low- $\nu_0$  technique



# Shape of $\nu_\mu$ or Anti- $\nu_\mu$ Flux using Low- $\nu_0$ Method

$\nu_\mu$ , Low-Nu0 Fit, ND at 500m

Relative  $\nu_\mu$ -Flux Measurement using Low- $\nu_0$  @ LBNE



Conclusion ➡  
Predict FD/ND flux-ratio with high precision

## $\pi\pi^0$ -Reconstruction

🔍 Clean  $\pi\pi^0$ - and  $\gamma$ -signatures in HiResMnu(STT)

🔍  $\nu$ -NC & CC  $\Rightarrow \pi\pi^0 \Rightarrow \gamma\gamma$

~50% of the  $\gamma \Rightarrow e^+e^-$  will convert in the STT, away from the primary vertex. We focus on these

🔍  $\gamma$ -Identification:

✱  $e^-/e^+$  ID: TR

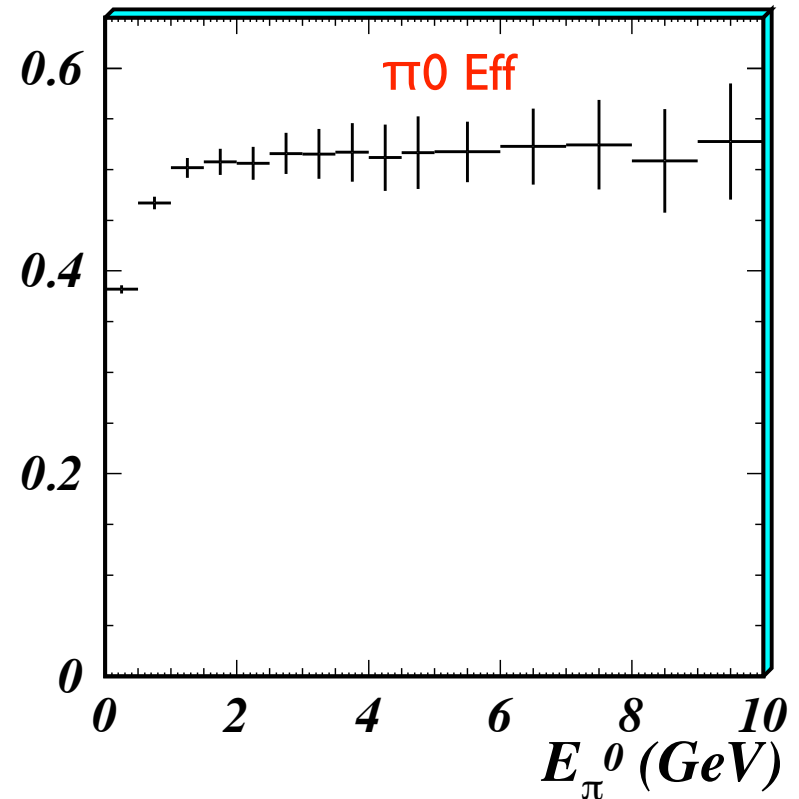
✱ Kinematic cut: Mass, Opening angle

➤ At least one converted  $\gamma$  in STT

(Reconstructed  $e^-$  &  $e^+$ ;  
 $e^-$  or  $e^+$  traverse  $\geq 6$  Mods)

➤ Another  $\gamma$  in the  
Downstream & Side ECAL

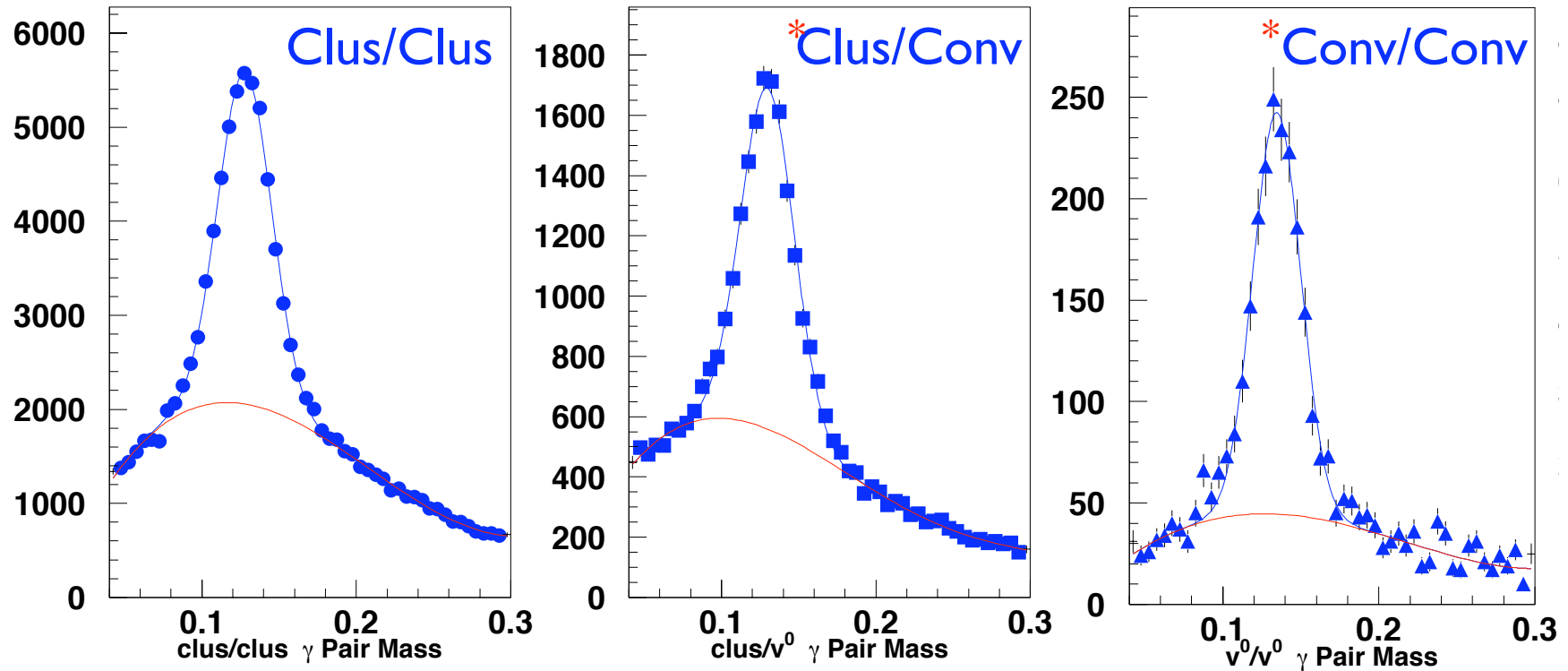
Efficiency



Conclusion  $\Rightarrow$

$\pi\pi^0$ 's Very well constrained in CC and NC

## Reconstructed $\pi^0$ in NC interactions in NOMAD



*Overall more than 33k reconstructed events. Three topologies:*

- *Cluster/Cluster 24k events*
- *Cluster/Conversion 7k events*
- *Conversion/Conversion 2k events*

[STT: expect similar resolution but  
much lower combinatorics]

South Carolina Group

## MEASUREMENT OF THE RATIO $\mathcal{R}_{e\mu}$ $\Leftarrow$ Search/Impact of Large- $\Delta m^{**2}$ Oscillation

- ◆ Independent analysis of neutrino data and anti-neutrino data due to possible differences following MiniBooNE/LSND results

$\Rightarrow$  Need a near detector which can identify  $e^+$  from  $e^-$

- ◆ Measure the ratio between the observed  $\nu_e(\bar{\nu}_e)$  CC events and the observed  $\nu_\mu(\bar{\nu}_\mu)$  CC events as a function of  $L/E_\nu$ :

$$\mathcal{R}_{e\mu}(L/(E\nu)) \equiv \frac{\# \text{ of } \nu_e N \rightarrow e^- X}{\# \text{ of } \nu_\mu N \rightarrow \mu^- X}(L/(E\nu))$$

$$\bar{\mathcal{R}}_{e\mu}(L/(E\nu)) \equiv \frac{\# \text{ of } \bar{\nu}_e N \rightarrow e^+ X}{\# \text{ of } \bar{\nu}_\mu N \rightarrow \mu^+ X}(L/(E\nu))$$

- ◆ Compare the measured ratios  $\mathcal{R}_{e\mu}(L/E\nu)$  and  $\bar{\mathcal{R}}_{e\mu}(L/E\nu)$  with the predictions from the low- $\nu_0$  flux determination assuming no oscillations  $\Leftarrow$  Benefit from External  $K^+/\pi^+, K^-/\pi^-, K^0_L/K^+$
- ◆ Same analysis technique used in NOMAD to search for  $\nu_\mu \rightarrow \nu_e$  oscillations.



ainment of the events so reducing the usable statistics.

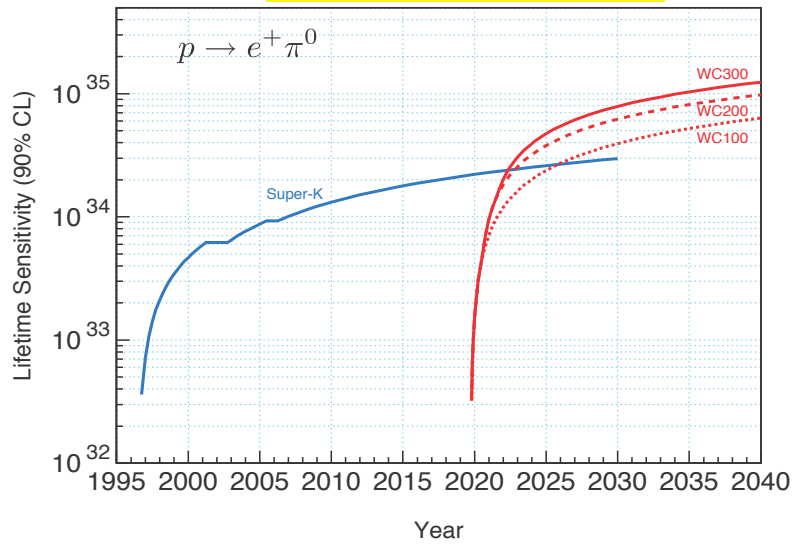
Measurement	STT	Sci+ $\mu$ Det	LAr	LArB	LArB+Sci+ $\mu$ Det	LAr+STT
In Situ Flux Measurements for LBL:						
$\nu e^- \rightarrow \nu e^-$	Yes	No	Yes	No	No	Yes
$\nu_\mu e^- \rightarrow \mu^- \nu_e$	Yes	Yes	No	Yes	Yes	Yes
$\nu_\mu n \rightarrow \mu^- p$ at $Q^2 = 0$	Yes	Yes	No	No	Yes	Yes
Low- $\nu_0$ method	Yes	Yes	No	Yes	Yes	Yes
$\nu_e$ and $\bar{\nu}_e$ CC	Yes	No	No	Yes	Yes	Yes
Background Measurements for LBL:						
NC cross sections	Yes	Yes	No	Yes	Yes	Yes
$\pi^0/\gamma$ in NC and CC	Yes	Yes	Yes	Yes	Yes	Yes
$\mu$ decays of $\pi^\pm, K^\pm$	Yes	No	No	Yes	Yes	Yes
(Semi)-Exclusive processes	Yes	Yes	Yes	Yes	Yes	Yes
Precision Measurements of Neutrino Interactions:						
$\sin^2 \theta_W$ $\nu$ N DIS	Yes	No	No	No	No	Yes
$\sin^2 \theta_W$ $\nu e$	Yes	No	Yes	No	No	Yes
$\Delta s$	Yes	Yes	Yes	Yes	Yes	Yes
$\nu$ MSM neutral leptons	Yes	Yes	Yes	Yes	Yes	Yes
High $\Delta m^2$ oscillations	Yes	No	No	Yes	Yes	Yes
Adler sum rule	Yes	No	No	No	No	Yes
$D/(p+n)$	Yes	No	No	No	No	Yes
Nucleon structure	Yes	Yes	Yes	Yes	Yes	Yes
Nuclear effects	Yes	Yes	Yes	Yes	Yes	Yes

TABLE XXVIII: Summary of measurements that can be performed by different ND reference configurations.

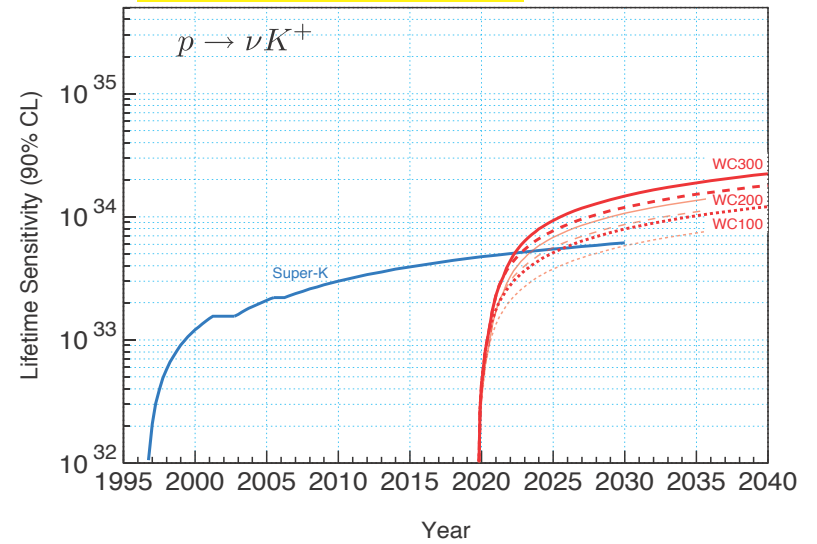
Summary page from the Short-Baseline Physics Report:

# Proton Decay

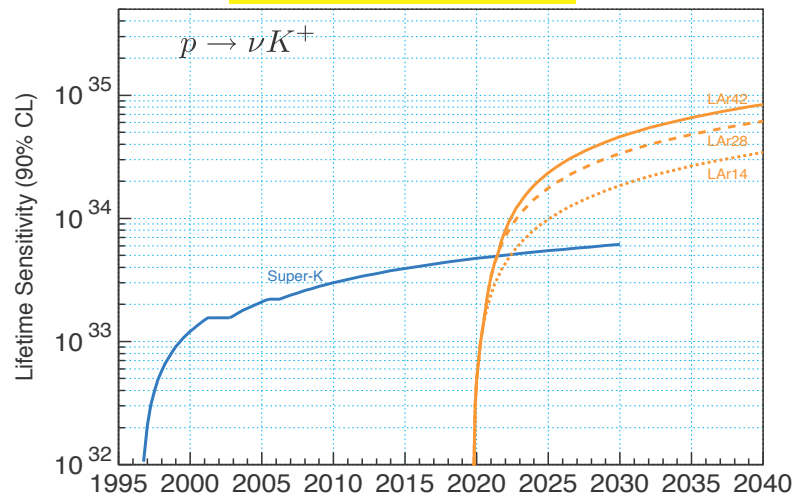
$$p \rightarrow e^+ \pi^0 \text{ WCD}$$



$$p \rightarrow K^+ \nu \text{ WCD}$$



$$p \rightarrow K^+ \nu \text{ LAr}$$

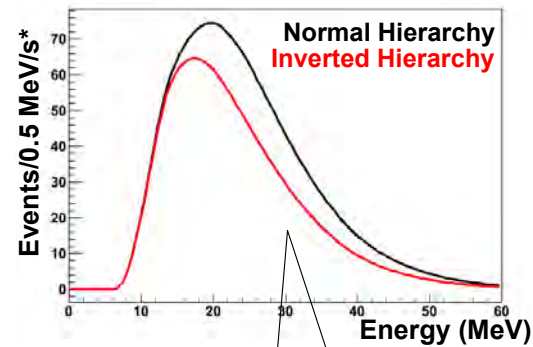


# SuperNova Burst

## Observed features in water and argon

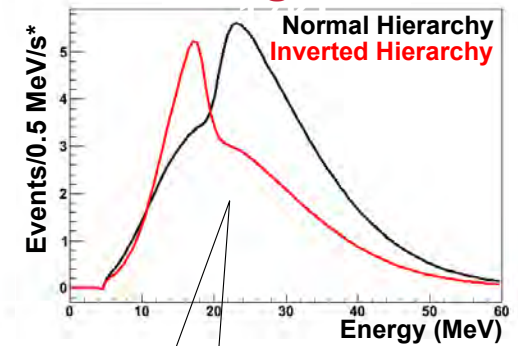


### Water



Larger detector,  
more events,  
but smooth

### Argon



Sharper,  
nonthermal  
features

Features in energy, flavour, time evolution: Depend upon progenitor, model, SN-type, Osci, ... ➤

\* one-second late-time slice

D. M. Webber

Channel	LAr 17kt Events
$\nu_e + {}^{40}\text{Ar} \rightarrow e^- + {}^{40}\text{K}^*$	1154
$\bar{\nu}_e + {}^{40}\text{Ar} \rightarrow e^+ + {}^{40}\text{Cl}^*$	97
$\nu_x + e^- \rightarrow \nu_x + e^-$	148
<b>Total</b>	<b>1397</b>

Channel	WCD 100kt Events
$\bar{\nu}_e + p \rightarrow e^+ + n$	27116
$\nu_x + e^- \rightarrow \nu_x + e^-$	868
$\nu_e + {}^{16}\text{O} \rightarrow e^- + {}^{16}\text{F}$	88
$\bar{\nu}_e + {}^{16}\text{O} \rightarrow e^+ + {}^{16}\text{N}$	700
$\nu_x + {}^{16}\text{O} \rightarrow \nu_x + {}^{16}\text{O}^*$	513
<b>Total</b>	<b>29284</b>

## Outlook

- 🐦 An comprehensive & ambitious  $\nu$  program at Fermilab leading to **LBNE**
- 🐦 Combination of  $L$ (Distance), **Beam**(1st & 2nd. Max), **Far** and **Near** Detectors
  - ⇒ Unique reach for the **Mass Hierarchy** and  $\delta_{CP}$ , possibly, something new!
  - ⇒ Sensitivity to **SuperNova & Relic- $\nu$** , Proton-decay, Atmospheric- $\nu$ , ...
- 🐦 The LBNE-ND aims to provide precise constraints on the systematic errors affecting the  $\nu$  oscillation physics:
  - ⇒ Flux of all 4 species:  $\nu_e$ ,  $\nu_\mu$  & **Anti- $(\nu_e, \nu_\mu)$**
  - ⇒ **Absolute  $E_\nu$ -scale**
  - ⇒ **Measurement of  $\pi^0$ /+/- --- backgrounds to oscillation-signal --- in NC and CC**
  - ⇒ **Difference between  $\nu$  & **Anti- $(\nu)$  interactions****
  - ⇒ *Provide a composite set of measurements which serve as the 'Event Generator' for the Far Detector*
- 🐦 A rich short-baseline  $\nu$ -physics
- 🐦 We welcome, *and need*, new institutions/collaborators

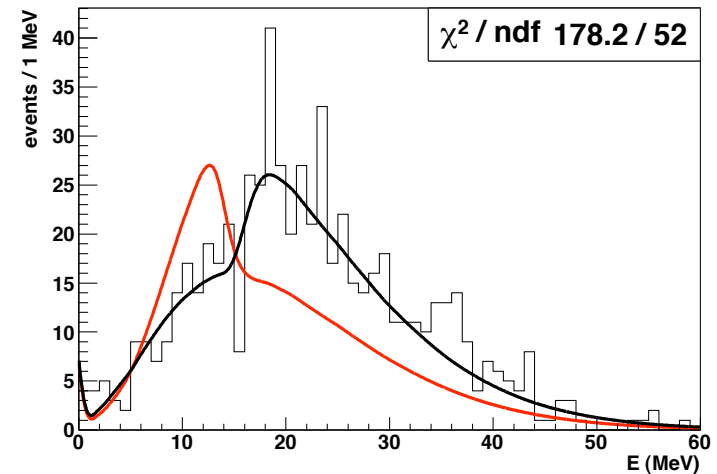
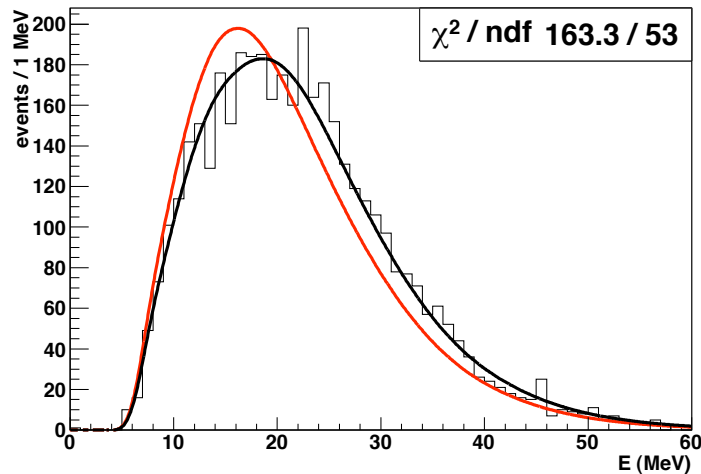
## *Backup Slides*

## SN $\nu$ -Spectra and MH

The SR neutrino spectra from 10 kpc observed in the LBNE detectors for normal hierarchy and the fit to the opposite hierarchy:

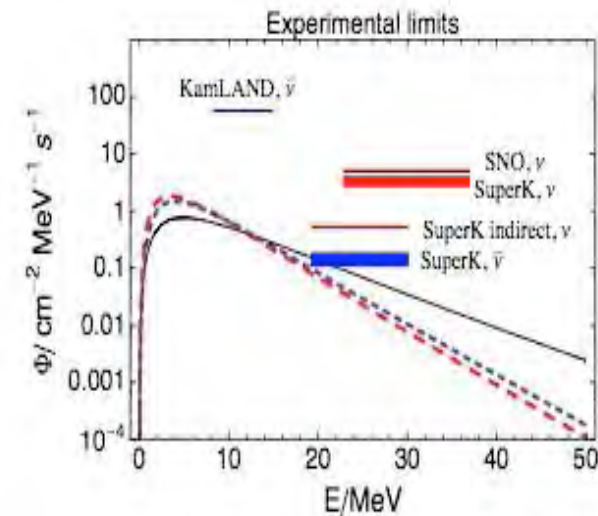
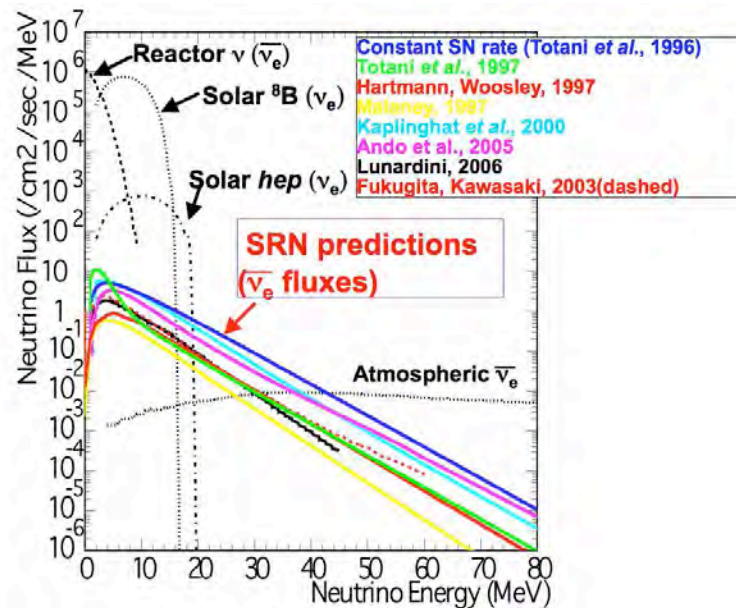
100 kt WCD - mostly  $\bar{\nu}_e$

17 kt LAr - mostly  $\nu_e$



Some ability to distinguish the MH - backgrounds yet to be evaluated

## Relic Supernova $\nu$



Reference Configuration	Expected Annual SRN Signal (events/year)	Expected Annual Background (events/year)	Years of LBNE Data Needed for a 3.0- $\sigma$ Signal Assuming Maximum/Minimum SRN Flux
300kt WCD 30%	5 – 52	320	1.3/144
300kt WCD 30% + Gd	13 – 74	64	0.13/0.9
100kt WCD + 100kt WCD-Gd + 17kt LAr	5 – 39	114	0.35/3
100kt WCD-Gd + 34 kt LAr	4 – 27	21	0.32/3

# Pionic correlations and meson-exchange currents in two-particle emission induced by electron scattering

J. E. Amaro,<sup>1</sup> C. Maieron,<sup>1</sup> M. B. Barbaro,<sup>2</sup> J. A. Caballero,<sup>1</sup> and T. W. Donnelly<sup>3</sup>

<sup>1</sup>Departamento de Física Atómica, Molecular y Nuclear, Universidad de Granada, E-Granada 18071, Spain

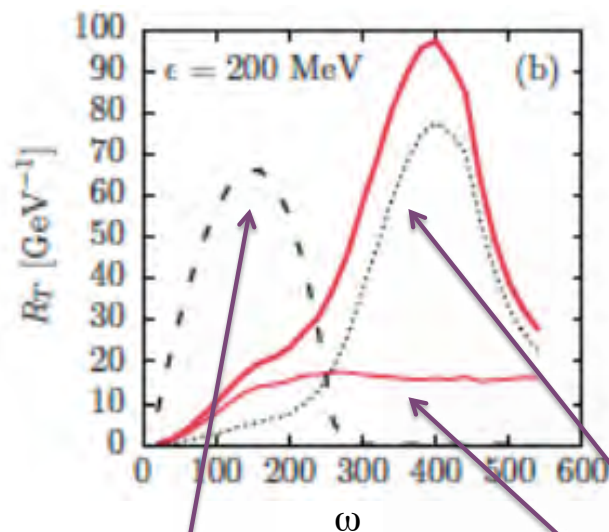
<sup>2</sup>Dipartimento di Fisica Teorica, Università di Torino and Istituto Nazionale di Fisica Nucleare, Sezione di Torino,

Via P. Giuria 1, I-10125 Torino, Italy

<sup>3</sup>Departamento de Física Atómica, Molecular y Nuclear, Universidad de Sevilla, Apdo. 1065, E-41080 Sevilla, Spain

<sup>4</sup>Center for Theoretical Physics, Laboratory for Nuclear Science and Department of Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

(Received 30 July 2010; published 4 October 2010)



One body RFG

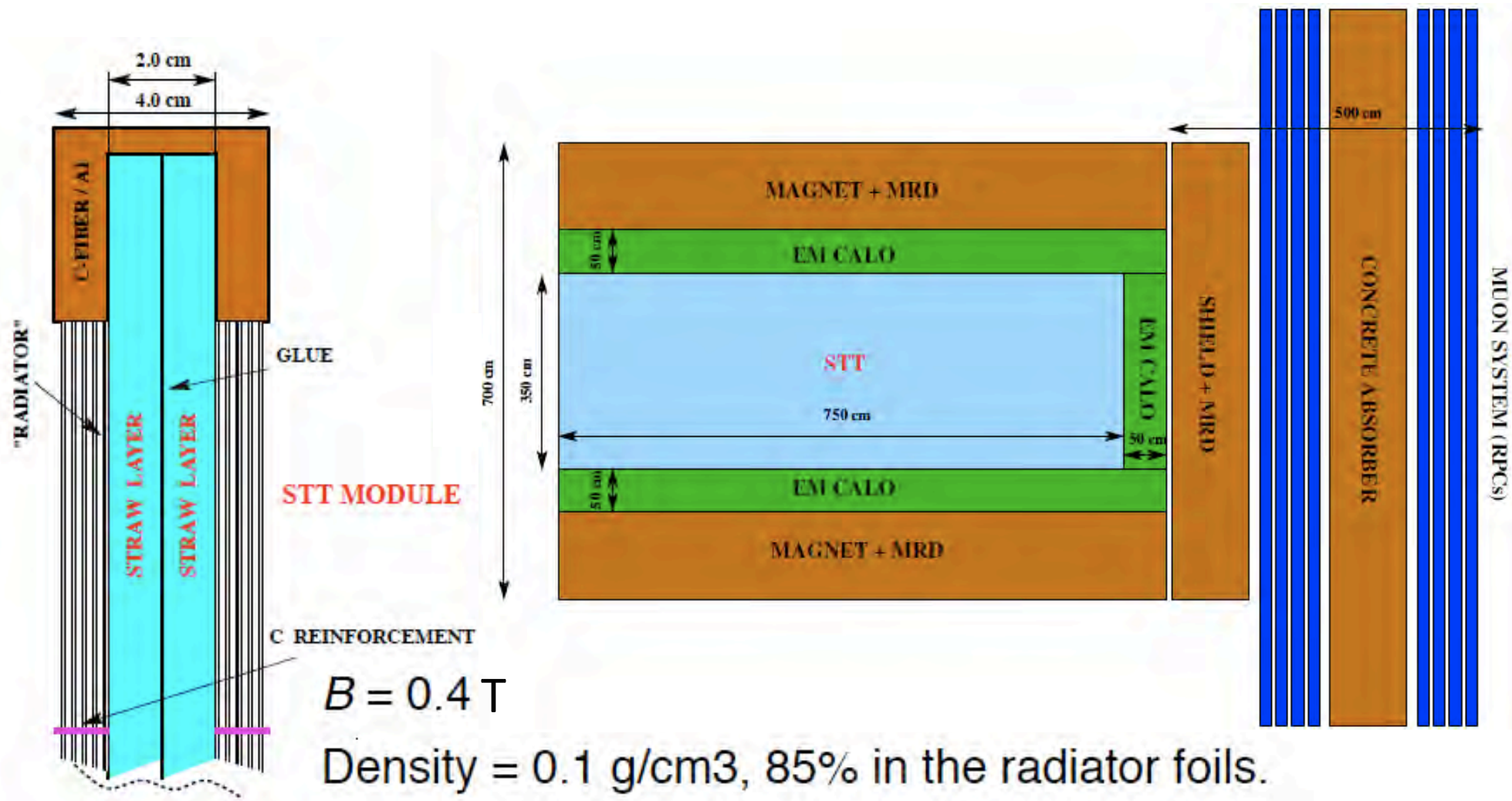
Meson-Exchange Currents predict a much larger fraction of the incident neutrino energy going into the hadron sector. Neutrino & antineutrino interactions may have different energy corrections up to ~300MeV and may create a spurious "CP-violating" effect, especially at 1.5 GeV where the sensitivity is maximum.

$R_T$  = transverse response function

Meson exchange

Correlation





Transition Radiation  $\Rightarrow e^-/e^+ \text{ ID} \Rightarrow \gamma$  (w. Kinematics)

$dE/dx$   $\Rightarrow$  Proton,  $\pi^+/-$ ,  $K^+/-$  ID

Magnet/Muon Detector  $\Rightarrow \mu^+/-$

## Resolutions in HiResMnu

•  $\rho \approx 0.1 \text{ g/cm}^3$

• Space point position  $\approx 200 \mu$

• Time resolution  $\approx 1 \text{ ns}$

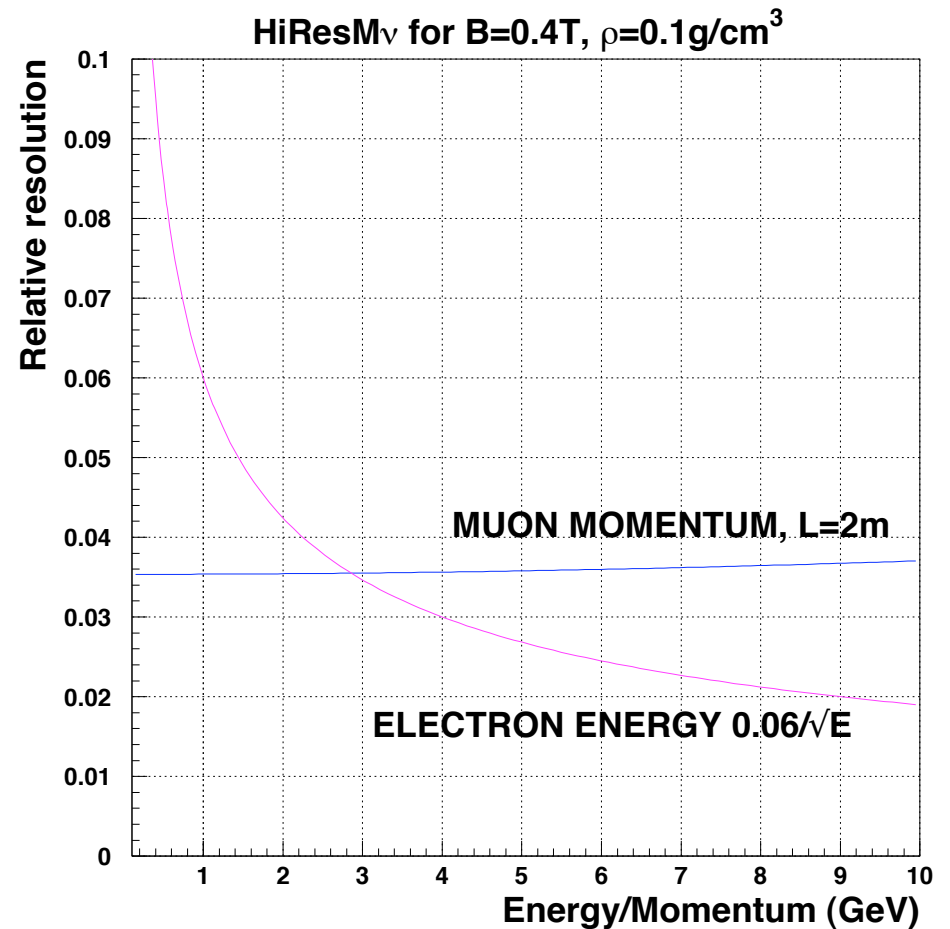
• CC-Events Vertex:  $\Delta(X,Y,Z) \approx O(100 \mu)$

• Energy in Downstream-ECAL  $\approx 6\%/\sqrt{E}$

•  $\mu$ -Angle resolution ( $\sim 5 \text{ GeV}$ )  $\approx O(1 \text{ mrad})$

•  $\mu$ -Energy resolution ( $\sim 3 \text{ GeV}$ )  $\sim 3.5\%$

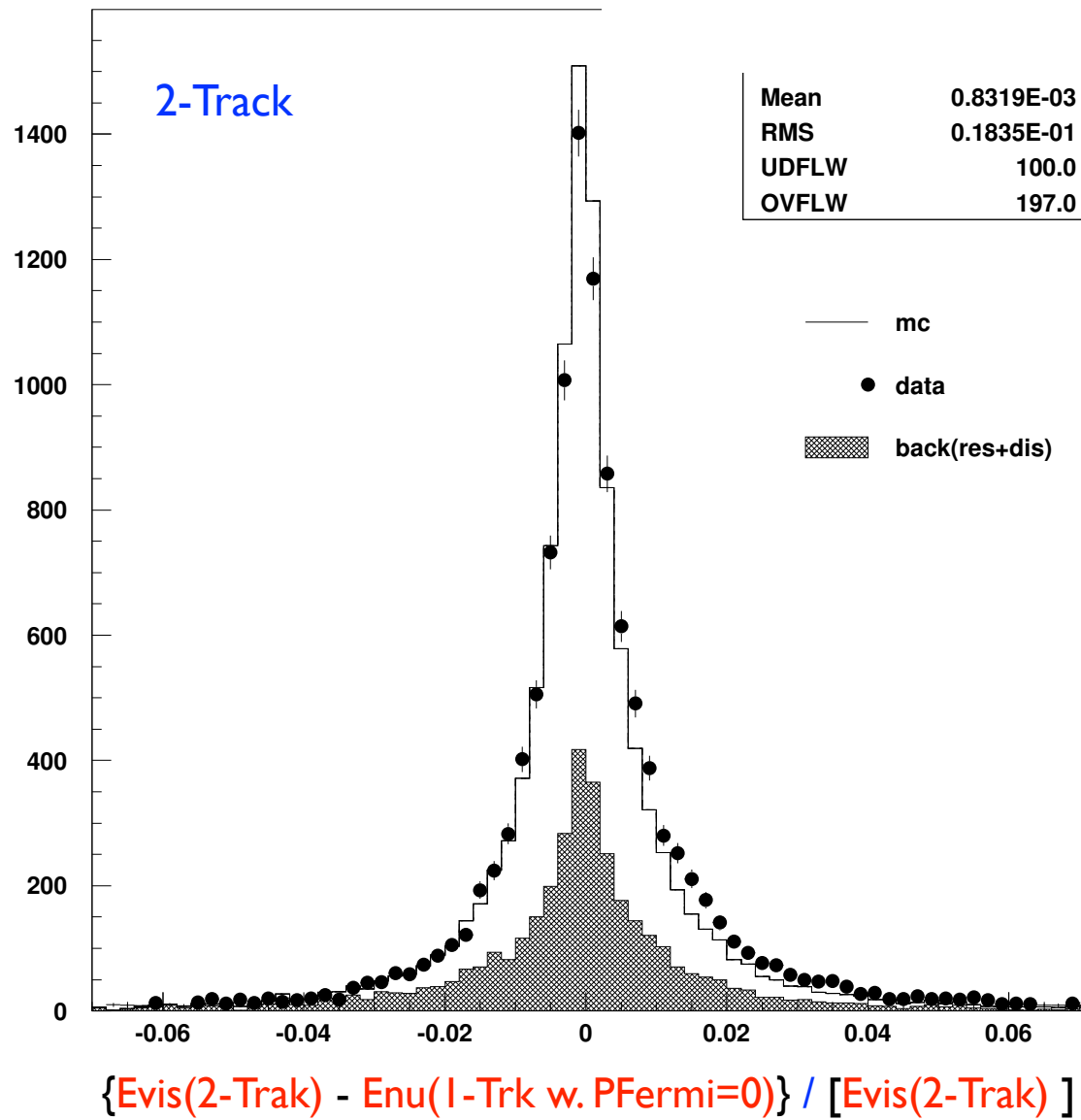
• e-Energy resolution ( $\sim 3 \text{ GeV}$ )  $\sim 3.5\%$



## *Sensitivity Calculations:*

• Parametrized calculation

• Repeat with NOMAD configuration and checked against the Data and Geant-MC (Agree within 15%): ReVt



$\Rightarrow$  constraint on  $E_\nu$  Scale

## Flux: ... Always the Flux

🙋 *Inverse Muon Decay*:  $\nu_x + e^- \rightarrow \nu_x + \mu^-$  {Single, forward  $\mu^-$ }

🐦  $\nu_\mu$  (t-channel) or Anti- $\nu_e$  (s-channel)

🐦 Elegant, Simple but steep threshold (calculable),  $E_\nu \geq 1.1 \text{ GeV}$

🐦 Systematic Advantage of STT lies in reducing systematic errors incurred by CCFR or CHARM-II in extrapolating the background to the signal  $\zeta = P_e(1 - \cos\theta_e) \leq \text{Cut}$

🙋  *$\nu$ -Electron Elastic Events*:  $\nu_x + e^- \rightarrow \nu_x + e^-$  {Single, forward  $e^-$ }

🐦 Different processes:  $\nu_e e^- \text{-CC}$ , Anti- $\nu_e e^- \text{-CC}$ , & all flavor  $\nu_x e^- \text{-NC}$

🐦 Different  $E_e$  spectrum

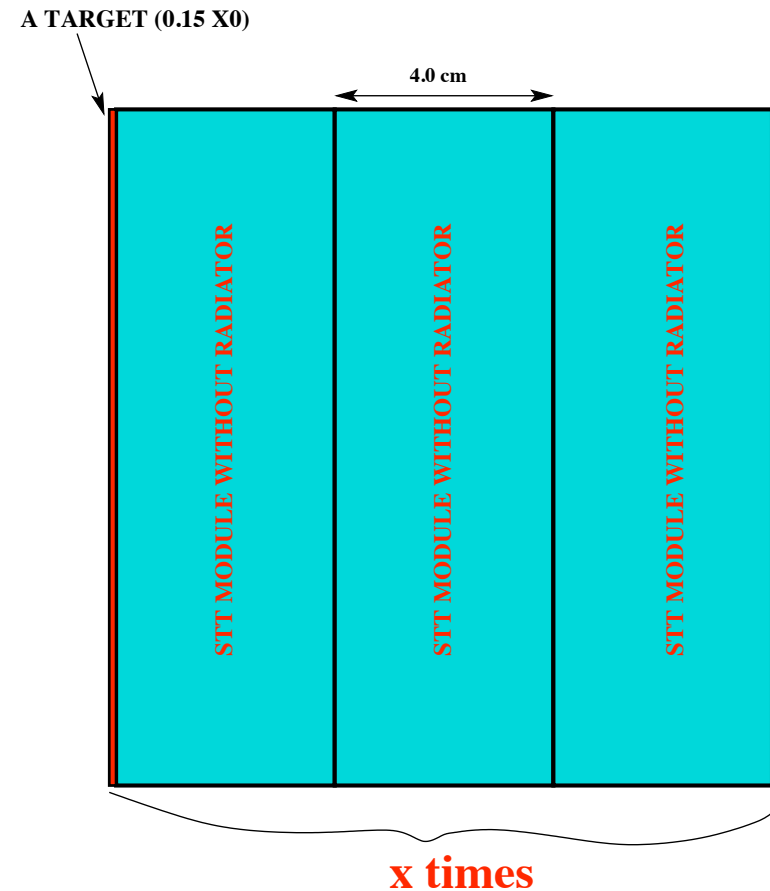
🐦 Focus on  $\nu_\mu e^- \text{-NC}$ : Experimentally the most challenging

☀ The Weak Mixing Angle (0.238) at  $Q^2 \sim 0.1 \text{ GeV}^2$  is known to  $\leq 1\%$  precision

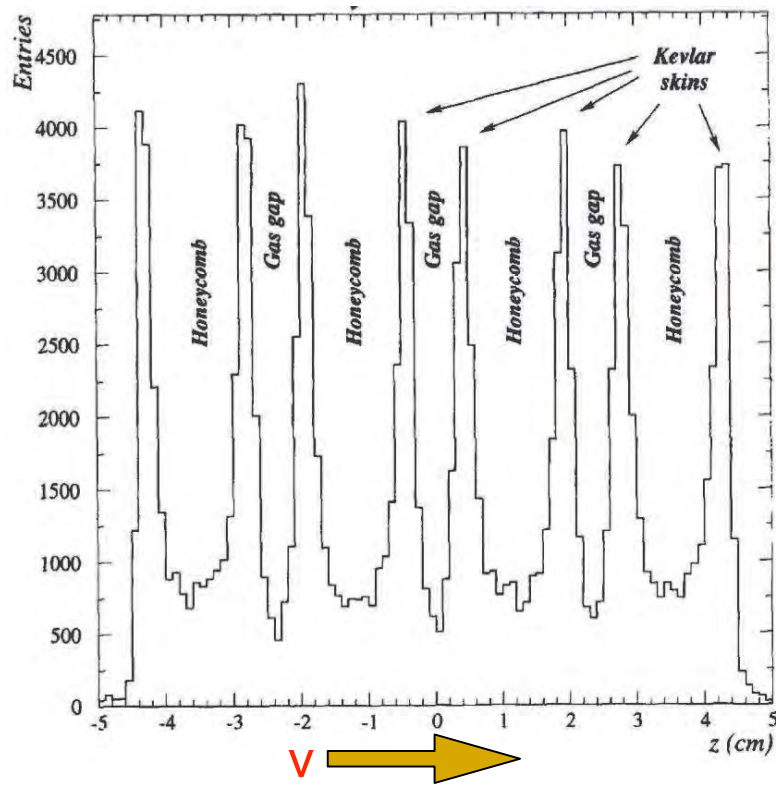
$\Rightarrow \sigma(\nu_x e^- \text{-NC}) \text{ known} \Rightarrow \text{Absolute-}\phi(\nu_x)$

## MEASURING NUCLEAR EFFECTS (Water, Ar, ..)

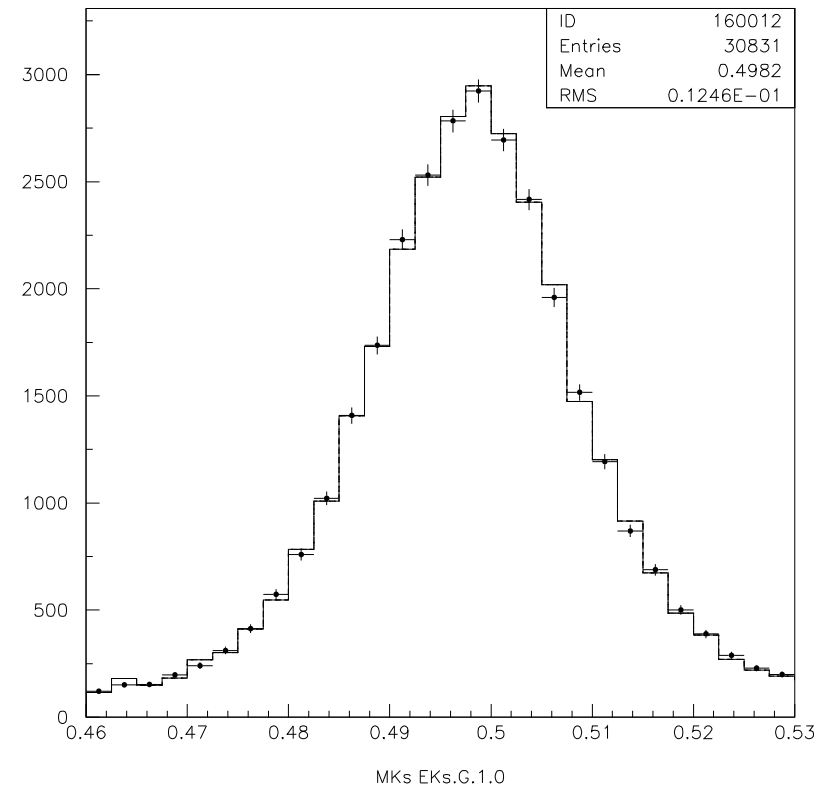
- ◆ Measure the  $A$  dependence (Ca, Cu,  $H_2O$ , etc.) in addition to the main  $C$  target in STT:
  - Ratios of  $F_2$  AND  $xF_3$  on different nuclei;
  - Comparisons with charged leptons.
- ◆ Use  $0.15X_0$  thick target plates in front of three straw modules (providing 6 space points) without radiators. Nuclear targets upstream.
  - For Ca target consider  $CaCO_3$  or other compounds;
  - **OPTION**: possible to install other materials (Pb, etc.).



## What we build on: NOMAD DATA



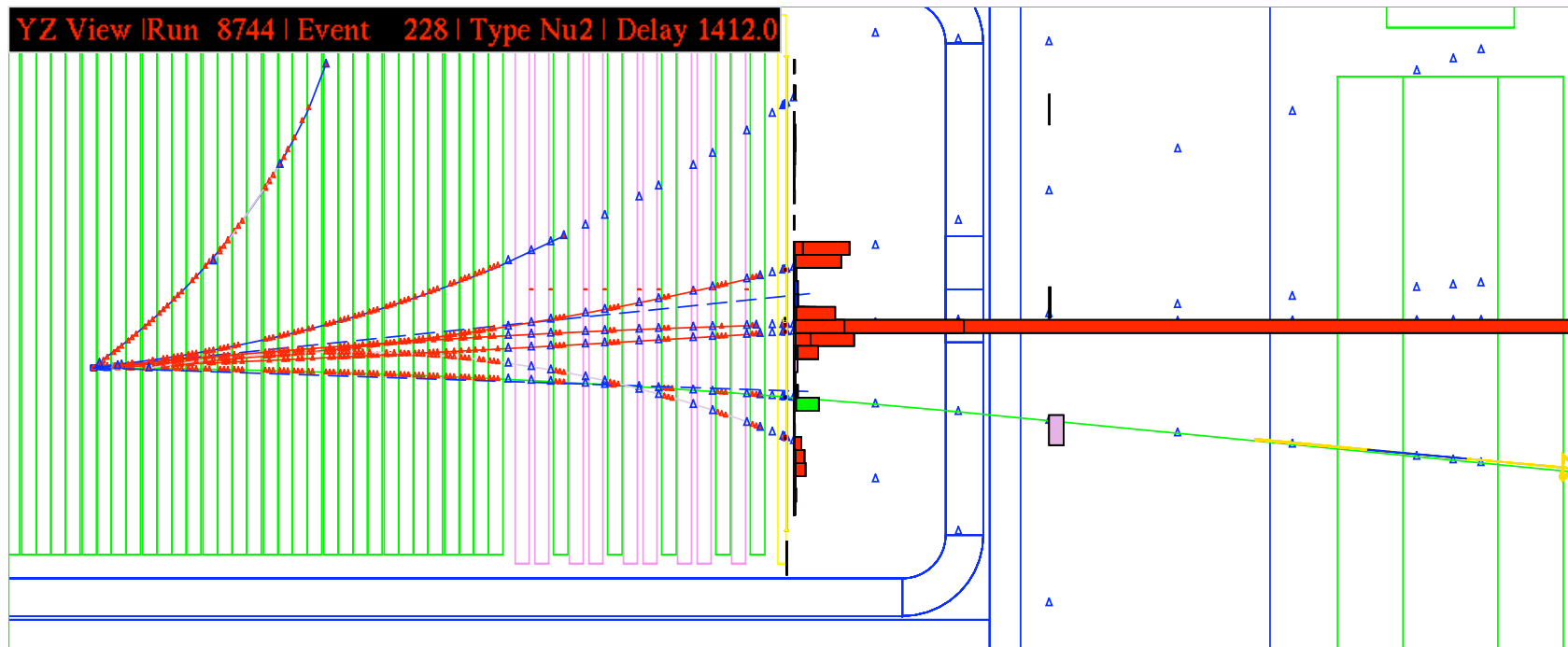
Neutrino radiography of one drift chamber



Reconstructed  $K^0$  mass

- ◆ NOMAD: charged track momentum scale known to  $< 0.2\%$   
hadronic energy scale known to  $< 0.5\%$
- ◆ HiResM $\nu$ :  $200 \times$  more statistics and  $12 \times$  higher segmentation

## A $\nu_\mu$ CC candidate in NOMAD



## LOW- $\nu_0$ METHOD

⇐ Shape of  $\nu_\mu$  or Anti- $\nu_\mu$  Flux

- ◆ *Relative flux vs. energy from low- $\nu_0$  method:*

$$N(E_\nu : E_{\text{HAD}} < \nu^0) = C\Phi(E_\nu)f\left(\frac{\nu^0}{E_\nu}\right)$$

*the correction factor  $f(\nu^0/E_\nu) \rightarrow 1$  for  $\nu^0 \rightarrow 0$ .*

⇒ *Need precise determination of the muon energy scale  
and good resolution at low  $\nu$  values*

- ◆ *Fit Near Detector  $\nu_\mu, \bar{\nu}_\mu$  spectra:*

- Trace secondaries through beam-elements, decay;
- Predict  $\nu_\mu, \bar{\nu}_\mu$  flux by folding experiental acceptance;
- Compare predicted to measured spectra ⇒  $\chi^2$  minimization

$$\frac{d^2\sigma}{dx_F dP_T^2} = f(x_F)g(P_T)h(x_F, P_T)$$

- *Functional form constraint allows flux prediction close to  $E_\nu \sim \nu^0$ .*
- ◆ *Add measurements of  $\pi^\pm/K^\pm$  ratios from hadro-production experiments to the empirical fit of the neutrino spectra in the Near Detector*

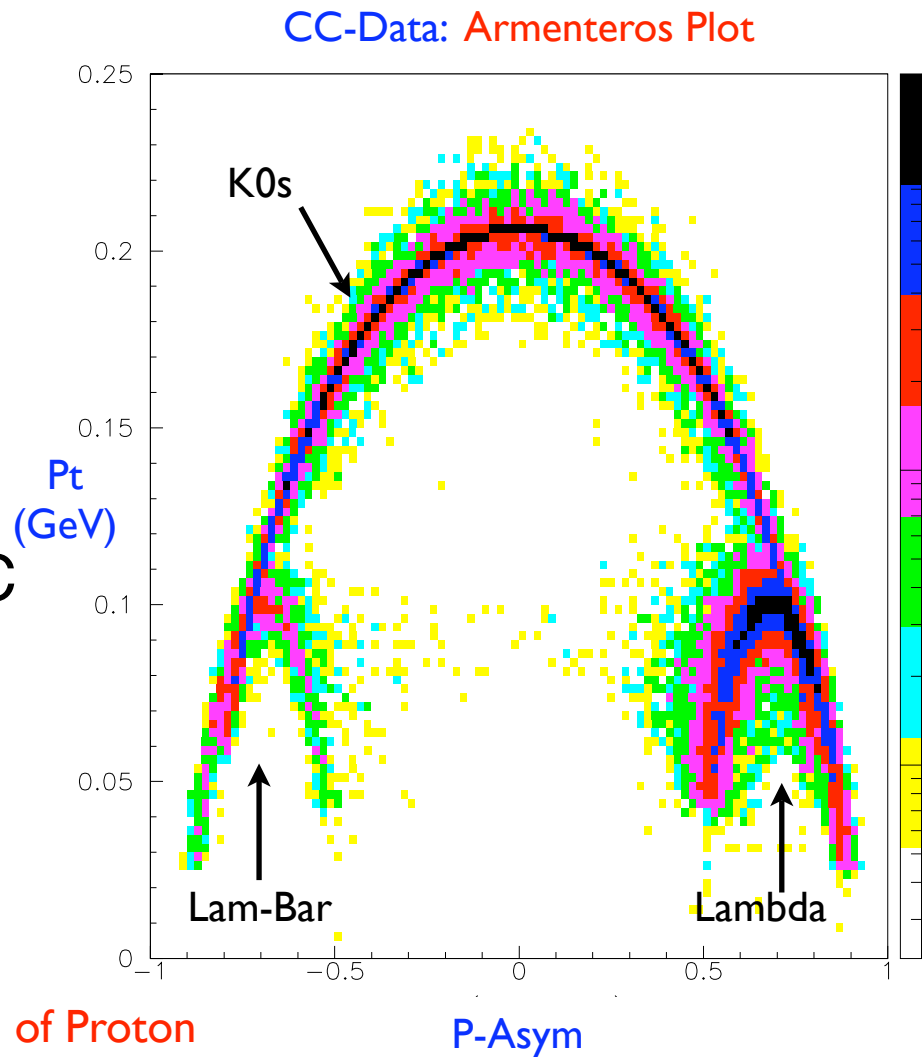


## Systematic-Errors in Low- $\nu_0$ Relative Flux: $\nu_\mu$ & Anti- $\nu_\mu$

- Variation in  $\nu_0$ -cut
  - Variation in  $\nu_0$ -correction
  - Systematic shift in Ehad-scale
    - Vary  $\sigma(\text{QE}) \pm 10\%$
    - Vary  $\sigma(\text{Res}) \pm 10\%$
    - Vary  $\sigma(\text{DIS}) \pm 10\%$
    - Vary functional-forms
  - Systematic shift in Emu-scale
  
  - Beam-Transport (ND at 1000m)
    - Includes:
      - \* Alignment (1.0mm)
      - \* Horn Current (0.5%)
      - \* Inert material (0.25 $\lambda$ )
      - \* Proton spot size
- ⇒ Revisit these (?) & Investigate ND @ 500m

# Measurement of exclusive topologies

- ◆ High resolution allows excellent reconstruction of exclusive decay modes
- ◆ NOMAD performed detailed analysis of strange particle production:  $\Lambda, \bar{\Lambda}$
- ◆  $\Delta$  resonances in CC & NC are easier to reconstruct
- ◆ Constraints on NC decay mode  $\Delta \rightarrow N\gamma$



$\Lambda$  Calibration of Proton  
Reconstruction